

NOTES TO THE CATALOGUE OF STARS.

The following notes to the stars include all those found in Dr. Peters' manuscripts. These consisted of brief notes and remarks all written in pencil on various papers. Some of his earlier notes, communicated to Harvard Annals, Vol. XIV, are superseded by later researches.

- 3. Long. Most authorities have $16^{\circ} 0'$, an error of $1 s = 16^{\circ}$, for $1 s' = 10^{\circ} 10'$.
Lat. Most Greek manuscripts have $74^{\circ} 20'$, and the Arabs $74^{\circ} 0'$ —either $\text{O}\Delta\Gamma'$ or $\text{O}\Delta$; it is more likely that the Γ' was omitted than that it was added.
- 6. Long. Paris 2389, Vat. 1594, and all the Arabs give $17^{\circ} 10'$. Manitius has $17^{\circ} 30'$.
- 12. Long. Baily gives $26^{\circ} 30'$.
- 13. Long. Baily and most Greek and Arab manuscripts have $26^{\circ} 40'$. Trapezuntius and Gerard of Cremona give $27^{\circ} 40'$, which has been adopted. Confusion in Arabic between 6 and 7 is very common, but it is not easy to explain an error in Greek of $s=6$ for $Z=7$.
- 18. Lat. Baily and all Greek manuscripts give $44^{\circ} 0'$. Sûfi, B. M. Reg. 16, and Bod. 369 have $45^{\circ} 0'$. All are clearly erroneous. Sûfi finds no fault with the position. The star is certainly Fl. 30 φ , which is described by Sûfi and was observed by Ulugh Beg. Peters conjectures that in the original uncial Greek $\Lambda\text{Z}\Delta' = 37^{\circ} 15'$ was written as shown in the Facsimiles (page 23) and thus resembled $\text{M}\Delta = 44^{\circ} 0'$.
- 25. Long. Baily gives $22^{\circ} 30'$.
- 26. Long. Baily gives $3^{\circ} 30'$.
- 37. Lat. All authorities agree. Latitude is 1° too large; it should be $\text{M}\Gamma' = 40^{\circ} 20'$, not $\text{M}\Lambda\Gamma' = 41^{\circ} 20'$.
- 41. Long. Baily gives $12^{\circ} 10'$. Nostar exists corresponding with the position in the Almagest. It was not identified by Baily or Schjellerup. Manitius considers it to be Fl. 8 Leo Minor. Peters conjectured that there was confusion in the Greek between $\text{I}\text{B}' = 12^{\circ} 10'$ and $\text{I}\epsilon\text{s}' = 15^{\circ} 10'$, which he adopts, and so arrives at the same star observed by Ulugh Beg (see photograph of Venice Codex 313, where ϵ in the abbreviation for $\text{M}\epsilon\text{i}\zeta\omega\text{v}$ might possibly be taken for β). Bod. 3374 has similar error of ϵ for β in the latitude, noted by Bernard about 1684. All the Arabs give latitude $22^{\circ} 45'$, Vat. 1594, $22^{\circ} 30'$.

	$\omega\psi$	s	Γ	$\omega\psi$	λS	Δ'
	$\omega\psi$	ϵ	$\Gamma\theta$	$\omega\psi$	$\lambda\Gamma$	Δ'
U.	$\omega\psi$	ϵ	Γ	$\omega\psi$	$\lambda\theta$	B
	$\omega\psi$	$\text{I}\epsilon\beta$	s	$\omega\psi$	$\mu\lambda\zeta$	δ'
	$\omega\psi$	Γ	s	$\omega\psi$	$\mu\lambda$	Γ
	$\omega\psi$	Γ		$\omega\psi$	$\mu\text{S}\zeta$	B
$\rho\upsilon\mu$	$\omega\psi$	$\text{I}\epsilon\beta$	$\Gamma\theta$	$\omega\psi$	$\text{I}\epsilon\theta\Gamma$	Γ
	$\omega\psi$	$\text{I}\epsilon\lambda$	s	$\omega\psi$	$\text{I}\epsilon\text{H}\lambda$	$\Gamma\theta$
	$\omega\psi$	λ	$\Gamma\theta$	$\omega\psi$	$\lambda\epsilon\lambda$	$\lambda\lambda$
	$\lambda\theta\psi$	$\theta\zeta$	Γ	$\omega\psi$	$\text{I}\epsilon\zeta\Gamma$	Γ
	$\lambda\theta\psi$	$\text{I}\Gamma$	Γ	$\omega\psi$	$\text{I}\epsilon\epsilon$	Γ
	$\lambda\theta\psi$	$\text{I}\beta$	s	$\omega\psi$	$\text{I}\epsilon\zeta$	B
	$\lambda\theta\psi$	IH		$\psi\psi$	$\text{H}\epsilon\Gamma\theta$	B
	$\lambda\theta\psi$	$\text{I}\epsilon\theta\zeta$	Γ	$\omega\psi$	$\text{H}\lambda$	B

FIG. 3.—Venice Codex 313.

PTOLEMY'S CATALOGUE OF STARS.

- 42. Lat. Most authorities give $23^{\circ} 0'$, but Ven. 313, Vat. 1594 and the Arabs have $20^{\circ} 20'$, which is right. Baily and Schjellerup could not identify. Manitius considers it to be Fl. 10 Leo minor. Peters finds that the star is VIII 245.
- 57. The large proper motion of 61 σ , $R. A. +0^{\circ}.0973$, $Dec. -1^{\circ}.766$, makes the identification of this star right.
- 58. Lat. Baily and all Greek manuscripts have $81^{\circ} 20'$; all the Arabs $81^{\circ} 40'$, which is adopted.
- 66. Lat. All Greek manuscripts have the incorrect latitude; the Arabs are right. $\text{P}\Gamma' = 80^{\circ} 20'$ for $\text{P}\Gamma = 83^{\circ} 0'$.
- 69. Long. All the Greek manuscripts have the erroneous longitude of $10^{\circ} 20'$; the Arabs are right. $\text{I}\Gamma'$ for $\text{I}\Gamma$.
- 75. Long. Baily adopts $5^{\circ} 10'$ from Gerard of Cremona. All the Greek manuscripts have the erroneous longitude of $9^{\circ} 0'$. The Arabs have $5^{\circ} 0'$, which is correct. One of the numerous errors of $\Theta = 9^{\circ} 0'$ for $\epsilon = 5^{\circ} 0'$.
- 79. Proper motion makes the disagreement in longitude much worse.
- 90. Long. Nearly all Greek manuscripts have the erroneous longitude $9^{\circ} 40'$. The Arabs are correct with $5^{\circ} 40'$. A similar error to No. 75.
- 96. This is the same star as No. 147.
- 97. Peters, Peirce, and Schjellerup identify this star as η Coronæ, which accords with the description, but the position agrees better with χ Bootis, adopted by Bode, Halma, Delambre, and Manitius. Baily is undecided between η and θ Coronæ.

Positions A. D. 100.

		η Coronæ.	χ Bootis.
Ptolemy Longitude....	$187^{\circ} 40'$	$190^{\circ} 20'$	$188^{\circ} 35'$
Latitude.....	$46^{\circ} 30'$	$47^{\circ} 1'$	$45^{\circ} 1'$

- 98. Not identified by Bode and Manitius. Baily and Schjellerup consider it to be χ Bootis, and Halma η Coronæ. The description accords best with Fl. 10 Coronæ.
- 99 to 102. There is much diversity of opinion as to the identification of these stars. Peters considered that they were in the following order: ω , b , ψ , and c ; Schjellerup as b , ω , ψ , and c ; Bode, Baily, and Manitius, c , ψ , b , and ω . The last accords best with the description and has been adopted. The comparisons for A. D. 100 are:

	Long.	Lat.		Long.	Lat.	Δl	Δb		Long.	Lat.	Δl	Δb
Ptolemy 12	$188^{\circ} 10'$	$41^{\circ} 40'$	ω	$187^{\circ} 10'$	$40^{\circ} 21'$	-60	-79	c	$188^{\circ} 33'$	$40^{\circ} 39'$	+23	-61
13	$186^{\circ} 40'$	$41^{\circ} 40'$	b	$188^{\circ} 19'$	$42^{\circ} 1'$	+99	+21	ψ	$186^{\circ} 59'$	$42^{\circ} 30'$	+19	+50
14	$187^{\circ} 0'$	$42^{\circ} 30'$	ψ	$186^{\circ} 59'$	$42^{\circ} 30'$	-1	0	b	$188^{\circ} 19'$	$42^{\circ} 1'$	+79	-29
15	$187^{\circ} 40'$	$40^{\circ} 20'$	c	$188^{\circ} 33'$	$40^{\circ} 39'$	+55	+19	ω	$187^{\circ} 10'$	$40^{\circ} 21'$	-30	-1

- 112. Lat. Baily and all Greek manuscripts have $46^{\circ} 30'$; the Arabs have $46^{\circ} 10'$, which agrees best.
- 129. Baily has Long. $3^{\circ} 40'$, Lat. $53^{\circ} 0'$. He remarks that there is no authority for latitude $50^{\circ} 40'$ adopted by Halma, but reference to the Table of Collations shows that nearly all Greek manuscripts have that latitude. Peters adopts $53^{\circ} 10'$ as in Sûfi and B. M. Reg. 16.
- 131. Lat. Baily gives $56^{\circ} 30'$.

134. Lat. All the Greek manuscripts give $63^{\circ} 0'$, which is wrong; the Arabs have the correct latitude, $60^{\circ} 20'$. Error of $\Xi\Gamma$ for $\Xi\Gamma'$.
135. Ptolemy's place is largely in error.
- 138, 139. Ptolemy's errors here are very large, and it is singular that the errors of the positions of these stars in Ulugh Beg are about as large. The identification of the stars is probably correct, but differs from Baily and Manitius.
140. Lat. Baily gives $72^{\circ} 15'$. The latitude $72^{\circ} 0'$ of the Arabs is adopted.
141. Lat. All the Greeks give $64^{\circ} 0'$, and the Arabs $60^{\circ} 15'$, which is correct. An error of $\Xi\Delta$ for $\Xi\Delta'$.
146. Long. The longitude agrees closely with the computed position, but considering the large errors in Ptolemy's longitudes of the stars in Hercules, it is probably 1° too large.
147. This is the same star as No. 96.
148. The identification of this star is probably correct, but the longitude and latitude are largely in error and no explanation of the discrepancy is available from the numerous manuscripts examined. Ulugh Beg has the correct latitude.
154. Long. There is no authority for the longitude $2^{\circ} 40'$ assigned by Peters to this star. All the manuscripts give $1^{\circ} 40'$. The very numerous errors in Greek of $A=1$ for $\Delta=4$ would suggest that here the longitude should be $4^{\circ} 40'$, which agrees closely with the computed place; but seeing the large errors in longitude common to all the stars in Lyra, it is doubtful if this explanation is available.
156. Identified as Fl. $9 \nu^2$, which agrees a little better and also is brighter than $8 \nu^1$, which Baily has taken.
159. Lat. Baily gives the latitude $49^{\circ} 0'$.
164. Long. Baily gives $19^{\circ} 20'$.
175. Peters considers this star the combination of $43 \omega^1$ and $45 \omega^2$. All the Greek and Arabic manuscripts give the latitude as $63^{\circ} 45'$, though $64^{\circ} 45'$, adopted by Baily, agrees closer; Halley gives $64^{\circ} 50'$.
184. Baily, Bode, Peirce, and Peters agree that this is ι Cassiopeia. Sûfi remarks that it is in a straight line with the two preceding stars δ and ϵ , which proves the identification correct. The longitude is 4° in error. All Greek and Arab authorities agree in Long. $1^{\circ} 40'$. The only explanation is an error in the earliest manuscripts of $A=1^{\circ}$ for $\Delta=4^{\circ}$, of which there are numerous instances in the manuscripts under discussion. Upon this explanation the difference of the computed place would be $+78'$, harmonizing with the general errors of the longitudes in Cassiopeia. The latitude is 1° in error, which is less easy to explain.
206. Lat. Baily and the Greek authorities give $28^{\circ} 0'$, and the Arabs $28^{\circ} 15'$, which is adopted.
221. Lat. Vatican 1594, Laurentian 1, Venice 313, and Paris 2390 are alike in giving the latitude as $\lambda\alpha \iota\Gamma'$. It is not clear what this means, but probably the iota has been written by mistake for the sign for $\eta\mu\sigma\nu$; thus it would be $31^{\circ} 50'$, as in Paris 2389 and Bod. 3374.
223. Lat. All authorities, except B. M. Arabic 7475, have latitude $20^{\circ} 0'$; the latter has $22^{\circ} 0'$, which is more nearly correct. Ulugh Beg has $21^{\circ} 30'$.
230. This is the same star as No. 400.
231. Long. All authorities, except B. M. Arabic 7475, have $26^{\circ} 0'$, which is 1° too small. B. M. 7475 has $27^{\circ} 0'$, which is correct. See note to No. 13.
233. Long. Baily gives longitude $20^{\circ} 40'$, latitude $16^{\circ} 20'$. Most of the Greek manuscripts have $20^{\circ} 40'$, an error conjectured of $K\Gamma$ for $K\Gamma'$. Paris Cod. 2394, $23^{\circ} 0'$, which is adopted. Grynæus $20^{\circ} 20'$, error of $K\Gamma'$ for $K\Gamma$. For

- latitude there are the readings $16^{\circ} 20'$ and $10^{\circ} 20'$; the latter is adopted. Sûfi remarks upon the erroneous position of Ptolemy, and Ulugh Beg did not find the star. The nearest star to the position is Fl. 5, but this is only 6.7 mag. The largest star in the neighbourhood is Fl. 2 of 5.0 mag., identified by Manitius, but this gives the large errors of Long. $-56'$ and Lat. $-150'$.
235. Lat. All authorities have 27° , which is 1° too small.
236. Lat. The Greek authorities have $26^{\circ} 30'$ and the Arabs $26^{\circ} 45'$.
239. There is great discordance in the manuscripts as to the coördinates of this star. The identification by Baily, Peirce, and Peters as Fl. 10λ is probably correct. The Arabs have the correct longitude. The latitudes, as appear in the table, are very discordant. Peters considered the latitude as $23^{\circ} 30'$ or $23^{\circ} 50'$. Cod. Ven. Greek 311, B. M. Reg. 16, and the Laurentian Arabic 156 have $23^{\circ} 45'$. Bodleian Arabic 369 has $28^{\circ} 45'$, which by the common error in Arabic of $\tau=8$ for $\tau=3$ may well accord. Probably $23^{\circ} 45'$ is the best to adopt. Baily has latitude $33^{\circ} 50'$.
246. Long. Baily and the Greek manuscripts give $26^{\circ} 40'$, which is erroneous; the Arabs and one reading of Paris 2389 have $23^{\circ} 40'$, which is correct. Peters remarks that if the Greek longitude is right, the star might be the Nova 1604, but Ulugh Beg observed the star 40ξ . This identification is confirmed by Peters, Baily, and Manitius.
- 247 to 250. Peirce states that these stars present one of the greatest perplexities of the whole catalogue. On reference to the Table of Collations, it will be seen that the manuscript authorities are about equally divided as to the latitude being north or south. Paris Codex 2389 gives both, which indicates that it is a compilation from more than one manuscript. Grynæus gives 247 as *north*, and omits any designation to 248-250, and it is singular that these are the only omissions in his whole catalogue of designation of the latitude, probably from the conflicting evidence in the manuscripts he used. The only printed editions which give the latitude of all these stars as *south* are Copernicus and Clavius. Peirce has discussed these stars in H. A. Vol. IX, but he is in error in stating that Baily has altered the latitude of the 16th star, No. 249. Peters' investigation leaves little room for doubt of his correct identification of the stars, and of their latitudes being *south*. The longitude of 250 is largely in error.
250. Lat. Baily has $0^{\circ} 45'$, which is found only in Liechtenstein and B. M. 7475.
251. There is some uncertainty as to the identification of this star. All manuscripts agree in longitude and latitude. Schjellerup and Manitius identify as Fl. 58, which would make the longitude erroneous by 2° and the latitude 1° . Bode and Halma give Fl. 2 ϵ (Sagittarius). The nearest star to the position is Fl. 52 (adopted by Baily), which is 6.6 mag. It has been conjectured that the star may have been Nova 1604, the position of which for A. D. 100 is longitude $236^{\circ} 44'$, latitude $+2^{\circ} 2'$, a difference of 1° in each coördinate. Peters does not decide between 52 Ophiuchi and 2 Sagittarii.
255. The Arabs have the correct latitude $1^{\circ} 40'$; Baily has $1^{\circ} 50'$.
262. Sûfi calls this a double star, which is Fl. 71 and 72 together.
268. Long. Most authorities and Baily give $23^{\circ} 10'$, which is 1° too small. Paris 2389 gives $26^{\circ} 10'$, which is nearer the computed place but is discordant with the other longitudes as being too large.
274. Lat. All authorities, Greek and Arabic, have latitude $16^{\circ} 15'$. But there is no suitable star in latitude 16° . Baily states that Bode and Delambre give it as $13^{\circ} 15'$, but without authority. Bode, however, gives it as $13^{\circ} 0'$.

There is no doubt that $13^{\circ} 15'$ is taken from Halley's edition of the Catalogue (*Geographiæ Veteris Scriptores Græci Minores, 1712*) which is a copy in which the positions of the stars have been corrected by computation.* It is probable that the identification of the star as Fl. 3 ν Ophiuchi is correct, and Ulugh Beg certainly observed this star. The latitude should be $13^{\circ} 15'$, and so it has been adopted by Peters. No explanation of the error in Greek is available.

285. Lat. Baily has $37^{\circ} 40'$, but the Arabs have $38^{\circ} 40'$, which is adopted.
289. Bode, Halma, Delambre, Baily, and Manitius make this star Fl. 54 α . Peters remarks that Fl. 59 ξ is Ulugh Beg's star and probably that of Ptolemy, but the latitude is 1° too large; besides ξ is 1 magnitude brighter than α .
296. Long. Bod. 3374 and Ven. 302 have $50^{\circ} 50'$, error of ν for η .
299. Longitude is 2° too large and latitude 1° too large.
300. Long. All authorities give $21^{\circ} 10'$, which is 1° too large. The position of this and the preceding star in Ulugh Beg are quite erroneous. Peters has adopted $20^{\circ} 10'$.
- 304 and 309. Long. In these stars longitude is 1° too small.
305. Lat. Adopted from Grynæus and Paris 2394. Most authorities give $33^{\circ} 50'$, which Baily adopts.
308. Lat. Several Greek and Arab authorities have $34^{\circ} 0'$. Error of $\Lambda\Delta$ for $\Lambda\Delta'$.
329. Long. Baily gives $9^{\circ} 10'$.
332. Long. Comparison with Ulugh Beg seems to indicate an error of 1° too large in Ptolemy's longitude.
346. Lat. Vat. 1594, Ven. 310 and 313, and all the Arabs have the correct latitude.
356. Lat. All authorities have $32^{\circ} 30'$, which is 1° too large, which is confirmed by comparison with Ulugh Beg.
357. Peters confirms Peirce in identifying this star as Fl. 1 α .
360. Long. There appears to be no authority for $16^{\circ} 40'$ adopted by Baily.
368. Latitude appears to be 1° too large; Ulugh Beg has $3^{\circ} 12'$; all authorities give $4^{\circ} 30'$ or $4^{\circ} 50'$.
371. The position of 63 τ^2 Arietis agrees much better than 61 τ^1 , and was certainly the star observed by Ulugh Beg.
372. Lat. Baily has $1^{\circ} 30'$.
374. The position agrees well with Fl. 87 μ Ceti (see note to 716 and 717, Ptolemy's 5 and 6 Ceti). Schjellerup, following Bode, identifies both 374 and 717 as μ Ceti. The agreement of Ulugh Beg with Ptolemy is so good that there can be no doubt that they observed here μ Ceti, while 717 does not agree at all. Manitius identifies 374 as Fl. 38, but the position for A. D. 100 is discordant. Δ long. = $+70'$; Δ lat. = $+107'$.
375. Lat. Baily has $10^{\circ} 30'$.
382. Long. Baily has $24^{\circ} 20'$, but the Arabs have probably the more correct longitude, $24^{\circ} 40'$.
383. Long. All Greek manuscripts, except Ven. 311, have erroneously $21^{\circ} 20'$. An error of $K\Delta = 21^{\circ}$ for $K\Delta = 24^{\circ}$.
389. Both longitude and latitude about 1° too large. Vat. Reg. 90 and Manitius give longitude as $10^{\circ} 20'$, an error of $1\Gamma' = 10^{\circ} 20'$ for $1\Gamma = 13^{\circ}$.
392. Ptolemy probably observed θ^1 and θ^2 as one mass.
394. Longitude $11^{\circ} 50'$ is adopted from all the Arabs, one reading of Paris 2389 and Ven. 312. Baily has $12^{\circ} 50'$, also from a variant in Paris 2389.

*The only available information about Halley's edition is the following paragraph from the preface to the above work: "Quod vero hisce omnibus subjungere placuerit Ptolemæi Catalogum Fixarum Stellarum, alicui forsan mirum videatur, cum sit argumenti plane dissimilis, minime tamen dubito quin hoc mihi ignoscat, qui norit quot ab illis syderibus maculas abstersit, quantamque eis lucem affundit Cl. Hallejus; eandem scilicet, qua, Ptolemæo illa contemplante, enituerunt: cum diu in libris, tam Mss. quam editis, ob voces perturbatas numerosque confusos, illa cæli lumina crassis obvoluta fuissent tenebris."

395. Long. The Arabs give $17^{\circ} 10'$ and the Greeks $17^{\circ} 30'$, as adopted by Baily; the first is preferable. Latitude in Paris 2389 is erroneously $0^{\circ} 15'$; error of $\Delta' = 0^{\circ} 15'$ for $\Delta = 4^{\circ} 0'$.
399. Lat. All authorities have $4^{\circ} 0'$, which is wrong; error of $\Delta = 4^{\circ} 0'$ for $\Delta' = 0^{\circ} 15'$. B. M. 7475 makes the latitude *north*, all the others *south*. Latitude $+0^{\circ} 15'$ would give the best accordance.
400. This is the same star as No. 230.
402. Lat. The Arabic Bod. 369 and B. M. Reg. 16 are the only authorities which have the correct latitude $0^{\circ} 15'$; all others, including Sûfi, have $4^{\circ} 0'$. Sûfi remarks that "Ptolemy's latitude is false, as the latitude places the star *north* of the preceding star, whereas the description states that it is *south*." This shows that the manuscript of Ptolemy used by Sûfi had the same error as in No. 399 above, viz., $\Delta = 4^{\circ} 0'$ for $\Delta' = 0^{\circ} 15'$.
404. Paris 2390, and the two Venice codices, 310 and 313, give the latitude correctly *south*.
405. Long. All manuscripts agree in giving $8^{\circ} 0'$; Manitius has $8^{\circ} 30'$.
406. The identification of this star is not free from doubt. Baily and Halma considered it to be 42 ψ Tauri and this star was finally adopted by Peters, but he remarks that Ulugh Beg's position of Ptolemy's 27th star in Taurus agrees fairly with 41 Tauri, but badly with 42 ψ . Ptolemy's star is in better harmony with 41 Tauri if we could assume an error of 1° in the latitude. The errors for A. D. 100 are:

	Δ Long.	Δ Lat.	Mag.
41 Tauri....	+ 2	- 57	5.3
42 ψ	+ 25	+ 32	5.3

Baily adopts latitude $7^{\circ} 20'$.

410. Long. Baily has $2^{\circ} 20'$.
412. Peters considered that there was no doubt that this star is III 170 and not Fl. 18 as Baily has, which gives errors for A. D. 100 of Long. $-27'$, Lat. $-19'$, mag. 5.6. III 170 gives errors of Long. $+51'$, Lat. $+9'$, mag. 5.4. The star can not be Alcyone. Ptolemy describes it distinctly as $\mu\kappa\rho\delta\varsigma$ (small). Gerard of Cremona gives mag. 5; all other authorities mag. 4.
415. The longitude $24^{\circ} 0'$ is adopted from one reading in Paris 2389, Venice 303, 311, 312, and the Arabs. The difference with other manuscripts is the common confusion of the alpha and delta. Baily has $21^{\circ} 0'$.
418. Peters, Peirce, and Manitius identify as Fl. 129 observed by Ulugh Beg, but the star is rather small and the longitude is too small. Peirce suggests that it might be better to make 418 as Fl. 126, and to suppose that 417 had disappeared. The position of Fl. 126 for A. D. 100 would accord very well with Ptolemy's star No. 418, but the identifications adopted accord best with the description.
- 419 to 423. Sûfi remarks that the longitudes and latitudes of these stars are grossly in error. There seems little doubt that Peters' identification is correct. Ulugh Beg's positions agree fairly well with them. They are all small stars.
424. Lat. Baily and all the Greeks have $9^{\circ} 30'$, and the Arabs $9^{\circ} 40'$, which is more correct.
426. Long. All authorities agree, but the longitude is 2° too large. The latitude is too small. Bod. Arabic 369 gives $11^{\circ} 0'$, which is more nearly correct.
432. Baily adopts longitude $26^{\circ} 10'$, latitude $3^{\circ} 0'$. The Greek manuscripts give longitude $26^{\circ} 10'$, and the Arabs $23^{\circ} 10'$; the latter is certainly the better to

- adopt. The latitudes are either $\Gamma = 3^\circ 0'$ or $\Gamma' = 0^\circ 20'$. Adopting the latter, the position agrees with Fl. 58. Baily identifies as 76 c. Peirce as 52 Tauri, Schjellerup as b, and Manitius as 63.
434. Lat. Baily adopts $18^\circ 15'$ from all authorities. Peters gives the longitude as $18^\circ 10'$ for the reasons given on page 12 for believing that the instrument used for measuring longitudes was not graduated to $15'$.
436. Long. Baily has $21^\circ 20'$. There is great uncertainty in the latitude of this star in all Greek manuscripts and in the printed Greek of Grynæus and Halma. In all cases it is represented by the character for $\frac{1}{2}$ followed by that for $\frac{1}{8}$, or in Paris 2389 and Laurentian 1, by 6. There is a slight indication in Paris 2389 (though not in Laurentian 1) of a separation of 6 from $\frac{1}{2}$, in which case it may be possibly $\frac{1}{2}$ with 6° as a variant. Peters considered the majority of cases he examined to be $0^\circ 30'$ with variant $0^\circ 10'$, not $0^\circ 40'$. All the Arabs agree in latitude $6^\circ 0'$, which is adopted.
438. Long. The better reading is that given by the Arabs and Vienna 14.
- 445 and 446. Baily, who took the Greek descriptions of the stars from Grynæus, did not perceive the error in the descriptions of these two stars, which are equally erroneous in Paris 2389. He gives:
445. τῶν ἐπομένων τῇ δεξιᾷ χειρὶ τοῦ ἐπομένου διδύμου ὁ μέσος τῶν γ. (τριῶν).
446. ἐπ' εὐθείας ὁ βόρειος. It is obvious that these descriptions should be as in Vatican 1594 thus:
445. τῶν ἐπομένων τῇ δεξιᾷ χειρὶ τοῦ ἐπομένου διδύμου τριῶν ἐπ' εὐθείας ὁ βόρειος.
446. ὁ μέσος τῶν τριῶν. Baily also states that the latitude of 445 in Paris 2389 is $-2^\circ 40'$, but in that manuscript it is clearly $-1^\circ 20'$.
- 445 to 448. The longitudes of these stars are all in error. The authorities give longitude of 448 as $0^\circ 40'$, except Laurentian 39, Vienna 14, and Vatican Reg. 90, which give $3^\circ 0'$, and Gerard of Cremona, B. M. Sloane 2795, which gives $5^\circ 40'$, the same as Liechtenstein; the last has been adopted. Peters remarks, "There is no other star than ζ Cancri that suits the position," hence the longitude is 1° too large.
449. Lat. Baily has $0^\circ 20'$. The value $0^\circ 40'$ given by the Arabs has been adopted as agreeing better with the computed position, and also by comparison with Ulugh Beg.
455. Ptolemy's position is erroneous. Ulugh Beg is right.
457. Lat. Baily and all authorities give $7^\circ 30'$. The error in latitude is remarked on by Sûfi and must be very old. Peters has adopted $10^\circ 30'$ without authority.
458. Long. All authorities have $19^\circ 10'$ (adopted by Baily) or $19^\circ 40'$, except Bodleian Arabic 369, and B. M. Reg. 16, which have $15^\circ 10'$. Sûfi remarked the error in longitude. There is little doubt the Arabs are correct, and we have another instance of error in the Greek of $\Theta = 9$ for $\epsilon = 5$. Peters identifies the star as the combination of 62 o' and 63 o'. Sûfi and Ulugh Beg both observed o Cancri. Baily, Schjellerup, and Manitius consider the star to be π Cancri.
459. Sûfi speaks of the error in longitude, which is 2° too large.
- 460 and 461. The latitudes of these two stars are wrongly transposed in all the authorities.
472. Long. All authorities agree, still the longitude is 1° too large. Ulugh Beg also has the longitude too large.
479. Long. Baily gives $12^\circ 10'$.
482. The identification of this star is one of the most difficult in the catalogue. Ptolemy states that it is the northern of two stars, the southern, No. 483, being well identified as θ Leonis. Fl. 81 is possibly the star, in which case Ptolemy's

- latitude would agree, but the longitude would be 4° in error. Peters remarks, "if we will not assume that a star disappeared near X 251, mag. 6.8, then the correction of longitude $1\text{H}\Gamma' = 18^\circ 20'$ for $1\Delta\Gamma' = 14^\circ 20'$ is the most plausible conjecture that can be made." There is, however, no evidence in the uncial Greek of papyri or of vellum manuscripts, nor in cursive Greek, of a confusion between $\text{H} = 8$ and $\Delta = 4$. "Sûfi speaks of the error in latitude of Ptolemy, but this can not be Ptolemy's star, and Sûfi had another star in view, while Ulugh Beg in his observations was guided by Sûfi." "Baily's identification with 71 Leonis is entirely to be rejected, since Baily himself has shown that the R. A. of 71 Leonis in Flamsteed by mistake is 2° too small."
486. Long. The authorities have either $24^\circ 40'$ or $21^\circ 40'$. The former is adopted, the latter is an error of $A = 1$ for $\Delta = 4$. The star is identified as 84 τ . Ulugh Beg observed 69 p'. Sûfi's description points to 74 φ .
487. Lat. All the Greek manuscripts, with the exception of Vat. Reg. 90, give the latitude as $3^\circ 12'$, which is clearly erroneous. There is no other instance in the whole catalogue of the fraction $\frac{1}{8}$. The error is doubtless of very ancient date. The magnitude of the star is $\epsilon' = 5$, and the latitude and magnitude are written thus: $\Gamma\epsilon'\epsilon'$. It is probable that in an early manuscript the magnitude was written by mistake within the latitude column, whence the mistake arose. Manitius has latitude $3^\circ 10'$ as Vat. Reg. 90. The Arabs have either $3^\circ 0'$ or $0^\circ 20'$, a confusion of Γ and Γ' . Latitude $3^\circ 0'$ is correct and so no doubt it was given in the original Greek.
- 494 to 496. The identification of these stars seems correct, and these were the stars observed by Ulugh Beg. The large error they have in common makes it look as if they were determined either differentially or by some other observer. Thus may be explained also why they are called *ἀμαυρός*, while not smaller than many others.

	C - Pt.	
	$\Delta l'$	Δb
	$^\circ$	$'$
494.	+2 29	-1 35
495.	+2 45	-1 34
496.	+3 25	-1 24

The following are the several identifications of the stars:

	Peters.	Baily.	Bode.	Halma.	Schjellerup.	Manitius.
494.	15 c	c	c	15 c	15 c
495.	7 h	4 Comæ.	h	h	12	7
496.	23 k	21 Comæ.	g	g	21	23

494 is given of magnitude 5, and is described by Ptolemy as *λαμπρός*. In Paris 2389 and Vat. 1594 it is *λαμπρός ἀμαυρός*; in the Trapezuntius edition "splendida," and in Liechtenstein, "luminosa." Ptolemy designates as *λαμπρός*, six stars mag. 1, thirteen mag. 2, seven mag. 3, and eleven mag. 4. He does not apply the word to any other star so faint as 494. It seems probable that here is a variable star.*

*These three stars of the *informata* of Leo, and described by Ptolemy as in the figure *πλόκαμος*, are three of the 12 stars which he designates as *ἀμαυρός*, the others being Nos. 40 to 43, among the *informata* of Ursa major, 219, the last of the *informata* of Perseus, and 311 to 314, the four stars in Equuleus. It is difficult to conjecture why these stars should have been designated *ἀμαυρός* (obscure). The magnitudes range from 4.1 to 5.1, the mean magnitude being 4.7. The constellations Equuleus and *πλόκαμος* are not mentioned by Aratus, Eratosthenes, Manilius, or Hipparchus in his commentary on Aratus. But Geminus (circa B. C. 77), in his work *Γεμίνου ἐισαγωγή ἐν τα φαίνομενα*, in his enumeration of the constellations, includes the constellation *προτομή ἵππου*, *sectio equi*. "according to Hipparchus"; and he also includes *βερρενικὴ πλόκαμος*, *Coma Berenices*. (Petavius, Uranologion, p. 12.)

- 497, 498. The longitudes of these stars are interchanged in all the manuscripts. Baily has not corrected them. The longitude $25^{\circ} 20'$ he gives to 497 should be that of 498. All the Greeks have $25^{\circ} 20'$, and the Arabs $26^{\circ} 20'$, which is adopted.
504. Peters remarks that the stars Fl. 44, 46, and 48, Virginis, mags. 5.9, 6.1, and 6.5, are near together, which may explain the greater brightness, mag. 5, estimated by both Ptolemy and Sûfi. Combined mag. 5.0.
509. Lat. Greek authorities give $20^{\circ} 10'$, the Arabs $15^{\circ} 10'$. Ulugh Beg's latitude is $16^{\circ} 15'$. Peters has adopted $16^{\circ} 0'$ from Halma, who is copied by Baily, and he remarks that Halma gives no authority. It is clear that Halma took $16^{\circ} 0'$ from Halley. It is of course correct, but is not supported by any manuscript.
513. Long. This is 1° too small; all authorities agree.
515. Peters and Baily agree that Ptolemy's position indicates 68 *i*, and both remark that it is clear that this position can not form the south following corner of the quadrilateral Ptolemy speaks of. But it is evident that the position of Ptolemy's 20th star in Virgo (correctly identified as 86) is exactly in the south following corner of the quadrilateral formed by 74, 76, and 82. The descriptions of Nos. 515 and 516 should be therefore interchanged.
517. Ptolemy's longitude is 2° too small, and the latitude error is similar in Ulugh Beg. This casts much doubt upon the identification of the star as 90 *p*, which, however, is not discordant with the description "in dextro crure posteriori." Peters questions whether there is here a variable or a star lost.
526. The identification as 53 is right, but Ptolemy's longitude is 2° too large. Ulugh Beg is also 1° too large. Baily gives latitude $7^{\circ} 10'$.
527. Ptolemy calls this star $\delta\pi\lambda\upsilon\upsilon\varsigma$; Sûfi likewise. The proper motion of Fl. 61 is so great, its distance from Fl. 63 ($73'$ in 1800) is reduced to $35'.4$ in Ptolemy's time. But could these two together appear double?
528. Peters agrees with Peirce in identifying this as 89, but the latitude is 1° too far south; Ulugh Beg likewise. Paris 2389, Vat. 1594, and the Arabs have the correct longitude, $5^{\circ} 0'$. Baily has $0^{\circ} 0'$.
529. The star is probably $\frac{a^1+a^2}{2}$ Libræ.
532. Long. Baily has $19^{\circ} 40'$, probably a misprint.
541. Ulugh Beg, misled by Sûfi, here probably observed 44 η , but Ptolemy's description does not admit this star. Greek authorities give latitude $3^{\circ} 0'$, which is probably an error of $\Gamma=3^{\circ} 0'$ for $\Gamma'=0^{\circ} 20'$. Bod. Arabic 369 and B. M. Reg. 16 have the latitude which has been adopted.
542. Peters identifies the position of this star as Oeltzen's Argelander, 14782, which has been found to be variable. Pickering remarks that it has not been observed brighter than mag. 9.
544. Lat. The Greek manuscripts have $8^{\circ} 30'$ and the Arabs $8^{\circ} 10'$, which latter is adopted. Baily has $8^{\circ} 30'$.
551. The star is $\frac{\omega^1+\omega^2}{2}$
553. This star, α Scorpii, is one of the six stars designated by Ptolemy as $\upsilon\pi\omicron\kappa\iota\acute{\rho}\rho\omicron\varsigma$; the others being α Bootis, α Tauri, β Geminorum, α Orionis, and α Canis Majoris. Questions relating to the color of these stars have been fully discussed by Nallino,* Schiaparelli,† Schjellerup,‡ and Knobel,§ including particular reference to the words used in Arabic texts as translation of the Greek. The word $\upsilon\pi\omicron\kappa\iota\acute{\rho}\rho\omicron\varsigma$ has been erroneously considered as

*Al Battāni. Pars II. †Rubra canicula. ‡Al Sûfi. Description des Etoiles Fixes.
§Monthly Notices, Vol. XLV; and Eleventh Oriental Congress, 1897.

- signifying *red*, its true meaning being "yellow, fire or wax-colored, cereus." and in that sense it has been correctly translated in the British Museum Arabic Almagest 7475, where the Greek word is expressed by the word شمدعي shemai, "wax-like;" but not so in Sûfi and all other Arabic texts. In these the Greek word is rendered by the sentence $\text{الموصي يضرِبُ الي الحوصي}$ or الموصي , meaning "inclines towards" some color expressed by الموصي . It is clear that this particular word is quite unknown to Arabs generally, and is not in any Arabic dictionary. All efforts to obtain a solution from scholars, and from the authorities at the Al Azhar Mosque at Cairo, have failed. Causin de Perceval,* speaking of another word used in Arabian Astronomy, says, "On chercherait en vain dans les dictionnaires Arabes et Latins l'explication de ce mot, et en général de presque tous les termes d'astronomie Arabe."† Ptolemy's designation of Sirius as $\upsilon\pi\omicron\kappa\iota\acute{\rho}\rho\omicron\varsigma$ has been exhaustively investigated by Schiaparelli and Schjellerup, who have shown the strong improbability of the term "Rubra Canicula" having been correctly applied to that star, or of there being any sound evidence of change in color. Though Sûfi omits all reference to the color of Sirius, yet in Bod. 369 and B. M. Reg. 16 the star is described by the same words indicating color as in the other five stars.
555. Lat. The Arabs $6^{\circ} 10'$ agrees a little better than the Greek $6^{\circ} 30'$ adopted by Baily.
560. Lat. All authorities, except Sûfi and Ulugh Beg, have $18^{\circ} 0'$; Sûfi $19^{\circ} 30'$. The star, according to Ptolemy's description, should be *south* of the preceding star and 18° does not agree at all; $19^{\circ} 0'$ has therefore been adopted.
567. Identified correctly by Peters as γ Telescopii. Ulugh Beg also observed this star. Ptolemy calls it nebulous. Peters says, "I can not see any nebulosity around it and Sûfi seems to doubt the same." There is, however, close to this star, the cluster N. G. C. 6441, described by Dreyer as "a globular cluster, very bright and pretty large." This seems to be the explanation of Ptolemy designating the object as nebulous.
569. The Greek authorities give the longitude as $25^{\circ} 30'$, which Baily has, and the Arabs $29^{\circ} 30'$, an error in the Greek of $\epsilon=5$ for $\theta=9$. The Greek latitude is $1^{\circ} 10'$, and the Arabs $4^{\circ} 10'$, a common error in Greek of $A=1$ for $\Delta=4$. In both elements the Arabs are right.
570. Long. All the Greek manuscripts give $9^{\circ} 30'$, except Ven. 312, $5^{\circ} 30'$ (same error as in the preceding note). The Arabs have $4^{\circ} 30'$, which is right. The confusion of θ or ϵ for $\Delta=4$ is not easily explicable.
577. The star is $\frac{\nu^1+\nu^2}{2}$. Ptolemy describes it as $\nu\epsilon\phi\epsilon\lambda\omicron\iota\delta\eta\varsigma$ και $\delta\pi\lambda\upsilon\upsilon\varsigma$. There are several small stars close.
578. Fl. 37 ξ^2 agrees better for position and is brighter than ξ^1 .
- 584 and 585. Ptolemy's large errors in longitude appear also in Ulugh Beg. Baily gives longitude of 584 as $25^{\circ} 20'$. There are no other stars corresponding.
587. Long. The Greek give longitude $22^{\circ} 40'$ and the Arabs $22^{\circ} 20'$, which latter is to be preferred. It is probable that 47 χ^1 and 49 χ^2 were observed as one mass.
592. Lat. The latitude is 1° too far south.
593. Long. Ptolemy's longitude is 2° too small. Sûfi remarks the error; Ulugh Beg is right.
596. Long. All the Greek authorities have $23^{\circ} 50'$ and the Arabs $26^{\circ} 50'$ —the latter is

*Notices et Extraits. Tome VII.
†M. D'Abbadie informed the writer that Fresnel told him that he learned in the Red Sea many current expressions not found in any native dictionaries.

adopted. Peters had $24^{\circ} 50'$ from Halley. The latitude in all the Greeks and some Arabs is $26^{\circ} 0'$. The only manuscript that gives the right latitude is B. M. Arabic 7475, $20^{\circ} 10'$. In the Greek there is an error of $K\epsilon$ for $K\varsigma$ '.

597. Long. Baily has $27^{\circ} 20'$.
604. Long. All authorities give the longitude either $9^{\circ} 0'$ or $5^{\circ} 0'$; similar error in the Greek, of which several examples have been given. Peters' adopted longitude of $6^{\circ} 0'$ is mere conjecture. It is more probable that the original was $5^{\circ} 0'$ and this was the opinion of Halley. Peters remarks that the proper motion of $2 \xi^2$ would bring the stars ξ^1 and ξ^2 quite close together in Ptolemy's time, only $5'.5$ apart, and that it was the combination of these stars that was observed.
609. As τ^2 is a little larger it was more likely to be the star observed, but perhaps $\frac{\tau^1 + \tau^2}{2}$ was observed as one mass.
610. Lat. Baily has $0^{\circ} 10'$ from Trapezuntius. The Arabs have $0^{\circ} 50'$, which is adopted.
- 611, 612, 613. See Baily's note on the confusion of these stars in different manuscripts. The description adopted agrees with Baily and Gerard of Cremona. Maniti-
tius adopts a different order.
613. Ptolemy's longitude is 1° too large.
615. Baily identifies as 35 Capricorni, mag. 6.0. Peters adopts 36 *b*, mag. 4.5, as being larger and more probable.
624. Liechtenstein and Sûfi erroneously designate the latitude *north*.
625. Ptolemy's longitude is too large.
626. Table of Collations shows that four Greek authorities (as well as Grynæus and Halma) have the erroneous longitude $20^{\circ} 40'$.
634. 13ν was the star observed by Ptolemy, whose longitude, however, needs a correction of $+2^{\circ}$.
635. The latitude appears to be 1° too small, though it agrees with Ulugh Beg.
642. Baily adopts latitude $2^{\circ} 10'$, which is erroneous.
645. Most of the authorities have latitude $4^{\circ} 0'$. Paris 2389 is correct; error of $\Delta = 4^{\circ} 0'$ for $\Delta' = 0^{\circ} 15'$. Peters identifies as 38 *e*, Baily as 37 *e*¹. Sûfi, misled by the erroneous latitude $4^{\circ} 0'$, observed Fl. 30. Maniti-
tius makes the latitude *south*.
649. Sûfi's observations point to 68 *g*² as the star which was observed by Ulugh Beg. Baily's identification as 59 *v* supposes an error of 3° in Ptolemy's longitude.
- 651 and 652. Peters identifies 651 as Fl. 63 *k*, but longitude and latitude are largely in error. The description of 651 is "Antecedens duarum quæ sunt in ipso aquæ fluxu a manu"; and the description of the following star, 652, is "Quæ istam adhuc sequitur." The latter star is correctly identified as Fl. 73 *l*. The star which precedes it and forms the pair referred to by Ptolemy is perhaps Fl. 67, though very uncertain, and it is smaller than 63 *k*. In the case of 63 *k* we have errors, longitude $-115'$, latitude $+136'$, and for Fl. 67 the errors are longitude $-106'$, latitude $-51'$. Baily identifies 651 as 67 and adds that a correction of $+2^{\circ}$ should be made to the longitude. Schjellerup identifies as Fl. 67. Sûfi omits 651 altogether.
657. The position is equally good for either 93 ψ^2 or 95 ψ^3 . The first is the larger star.
658. The star is probably Fl. 94. Sûfi seems to have observed Fl. 97, which gives errors of longitude $-113'$, and latitude $-106'$, and is smaller than 94. Ulugh Beg observed 94. All authorities give longitude $20^{\circ} 50'$, which is 3° too large. Upon this assumption Peters adopts $17^{\circ} 50'$.
- 659, 660. Baily gives the longitude of 659 as $22^{\circ} 20'$. There is no doubt that Ptolemy and Ulugh Beg observed ω^1 and ω^2 . It is curious that Sûfi remarks that

near one of these stars there is a star of mag. 6, which makes it double. Peters says it can hardly be the variable R Aquarii, which is 1° distant. It is probable that Sûfi really observed R at its maximum. The positions of ω^2 and R for 1875 are:

	R. A.			Decl.	
	h	m	s	°	'
ω^2 Aquarii . . .	23	36	14	-15	14.1
R Aquarii . . .	23	37	21	-15	58.6 var. mag. 6-11.

661. It is probable that the two stars A^1 and A^2 were observed as one mass $\frac{A^1 + A^2}{2}$.
663. Baily identifies as 106 *i*¹, but 108 *i*² agrees better; it is also described by Sûfi. Ulugh Beg seems to have observed 107.
- 665 and 666. The longitudes and latitudes are transposed in nearly all the manuscripts.
667. Lat. Peters' latitude, $16^{\circ} 15'$, is a conjecture; there is no authority for it, and there is no ready explanation of confusion in the Greek letters for $14^{\circ} 45'$ or $14^{\circ} 50'$ and $16^{\circ} 15'$.
668. Long. Baily has $12^{\circ} 20'$.
670. This is the same star as No. 1011.
685. Long. Baily has $20^{\circ} 10'$.
687. Longitude 1° too large, latitude 1° too far south.
688. Longitude adopted from Paris 2389, one reading, and Arabs. Baily has $23^{\circ} 20'$. Latitude 1° too far south.
689. Here Ulugh Beg has the south latitude too small.
690. Longitude of the Arabs adopted as more correct. Baily has $28^{\circ} 20'$.
694. Lat. Baily has $1^{\circ} 45'$, which is found only in Trapezuntius, Schreckenfuchs, and the Crawford manuscript of Gerard of Cremona.
695. Longitude of Arabs $0^{\circ} 20'$ is better than the Greek $0^{\circ} 40'$, which Baily adopts.
696. Peters identifies as the combination of 93, mag. 5.3, and 94, mag. 5.6, and adds that these two stars viewed as one mass would appear about mag. 4.7, so that the mean differences should be taken.
- 702 to 704. These are the stars observed by Ptolemy and described by Sûfi, but the positions are in error, as was noted by Sûfi. Maniti-
tius identifies 704 as χ , but though the position would suit, it is discordant with the description. Peters considered there was no doubt that No. 707 is correctly identified as χ , though the longitude is 2° too large.
- 716, 717. Baily gives the longitude of 716 as $10^{\circ} 20'$. These two stars present much difficulty. It is suggested that 716 may be either 78 ν , or 73 ξ^2 , but both give large errors in both elements. No star harmonizes with Ptolemy's position of 717. Schjellerup and Maniti-
tius identify as μ Ceti, but this star is more probably 374, Ptolemy's 13th star in Aries. The question of these two stars remains undecided.

	Ptolemy.		Position A. D. 100.	
	Long.	Lat.	Long.	Lat.
716.	$10^{\circ} 10'$	$-8^{\circ} 10'$	$11^{\circ} 58'$	$-9^{\circ} 21'$
717.	$12^{\circ} 40'$	$-6^{\circ} 20'$	$15^{\circ} 21'$	$-5^{\circ} 40'$

726. The latitude $15^{\circ} 20'$ of the Arabs has been adopted in preference to $15^{\circ} 40'$ of the Greek, which Baily has.
- 728 to 731. The identification of these 4 stars seems correct; they accord with the description. Longitude and latitude of 729 are largely in error. Ulugh Beg's latitude also in error.

734. Lat. The Greek manuscripts all have $16^{\circ} 30'$, with the exception of one reading of Paris 2389, and Cod. Ven. 303, which are $13^{\circ} 30'$. Sûfi and the Arabs have $13^{\circ} 50'$, or $18^{\circ} 50'$, which are equivalent by the common error of $\tau=8$ and $\bar{\tau}=3$. Baily remarks upon the error of 3° in the Greek authorities. Ptolemy describes this star as *νεφελοειδής*, probably from it making with φ^1 and φ^2 Orionis a small cluster.
738. Ptolemy's longitude seems 1° too large.
739. Ptolemy's longitude is too small, also when compared with Ulugh Beg.
740. Peters' identification is right. Ptolemy calls it *διπλούς*, probably from LL 11748 and LL 11884 being near and south of ξ .
741. Ptolemy's longitude 1° too large.
- 742 and 743. As Gore has correctly pointed out, the description of these stars should be reversed.
748. Lat. Baily has $20^{\circ} 10'$.
752. Baily denotes this as 6 g. Peters identifies as 9 o². The same deviations in longitude and latitude are found here as in Ulugh Beg. Baily's star 6 g. does not agree at all.
763. Lat. The Greek $28^{\circ} 20'$, which Baily has; the Arabs $28^{\circ} 40'$, which is adopted.
767. Long. The Greek $26^{\circ} 30'$, except Vienna 14, the Arabs $26^{\circ} 10'$, adopted, but longitude still too large.
774. Long. All the authorities have $48^{\circ} 0'$, which is 1° too large, also in comparison with Ulugh Beg.
775. Long. Paris 2394 has $\iota\delta u'$ in which the "u" is an old cursive form of β , and in this manuscript it would signify $14^{\circ} 40'$. Grynæus has $\iota\delta \epsilon' = 14^{\circ} 12'$.
777. Long. Several Greek manuscripts give $16^{\circ} 0'$ for $10^{\circ} 10'$; error of 1s for 1s'. Baily has $18^{\circ} 20'$, for which there is no authority.
778. Lat. Halma has $25^{\circ} 20'$, which he has taken from Halley.
779. Baily has longitude $3^{\circ} 30'$, and latitude $28^{\circ} 30'$. Peirce considers the star to be 98 Heis. Peters agrees with Baily and Schjellerup in identifying as 40 o².
781. Lat. All authorities give $32^{\circ} 50'$. Halma gives $33^{\circ} 10'$, which he has taken from Halley.
786. It is not possible to decide whether the star is ρ^2 or ρ^3 . Ptolemy observed them as one mass.
787. Lat. The Greek authorities give $23^{\circ} 30'$, while the Arabs have $23^{\circ} 50'$. Halma alone has $24^{\circ} 30'$, taken from Halley, and Baily adopts it. Peters did not notice the extracts from Halley made by Halma and Baily, and which he had adopted. In the present case the reading of the Arabs is taken.
788. Flamsteed remarks that a star noted by Ptolemy as of the 4th magnitude, and which is the 17th of the constellation Eridanus in his catalogue, could not be found now. About the position of the star all editions agree; it is the same in all existing manuscripts, both Greek and Arabic, and was the same also in the manuscript used by Sûfi. Sûfi says of this star: "The 17th, which precedes the 16th, is the last of the four, and at the western extremity of the series, near the four stars situated on the breast of Cetus. It is of the smaller ones of the 5th mag., almost of the 6th, and there is between it and the nearest star of the four situated on the breast of Cetus, that is, the 10th of Cetus, less than one 'coudée.'" Bode takes the star to be σ Eridani (Bayer and Ideler likewise), but says that since Flamsteed it is wanting upon all star charts and in the sky. Manilius takes it to be η , and the preceding star ρ^2 . According to Ptolemy's difference with η Eridani, the star could be Heis 10, 6.7 mag. = W. B. 2^h 788. According to Sûfi's description, the star seems to be nearer to ϵ Ceti (moins d'une coudée) than to η Eridani. He puts the distance between ρ and η Eridani as one coudée.

The following table shows the comparison between Ptolemy, Ulugh Beg, and computed positions, for A. D. 100, assuming the star to be W. B. 2^h 788:

Ptolemy's star.	Ptolemy.			Ulugh Beg reduced.		Name.	Computed.	
	Mag.	Long.	Lat.	Long.	Lat.		Long.	Lat.
16 Eridani..	3	12 10	-23 30	12 37	-24 35	η Eridani....	12 14	-24 41
17 Eridani..	4	10 30	-23 15	11 35	-24 17	W. B. 2 ^h 788.	11 7	-24 56
10 Ceti.....	4	6 40	-25 10	7 46	-26 20	ϵ Ceti.....	6 48	-26 7

We get the differences:

	17- η Eridani.		17- ϵ Ceti.	
	Long.	Lat.	Long.	Lat.
Ptolemy....	-1 40	+15	+3 50	+1 55
Ulugh Beg...	-1 2	+18	+3 49	+2 3
Computed...	-1 7	-15	+4 19	+1 11

The star W. B. 2^h 788 is therefore the nearest. Ptolemy calls 17 Eridani of the 4th magnitude, but Sûfi of the 5th magnitude, *small*, almost the 6th. In Harvard R. Photometry η Eridani is 4.0 mag. and ϵ Ceti 5.0 mag. The Uranometria Argentina gives the magnitude of W. B. 2^h 788 as 6.4. In the following chart the position of W. B. 2^h 788 is marked by a +.

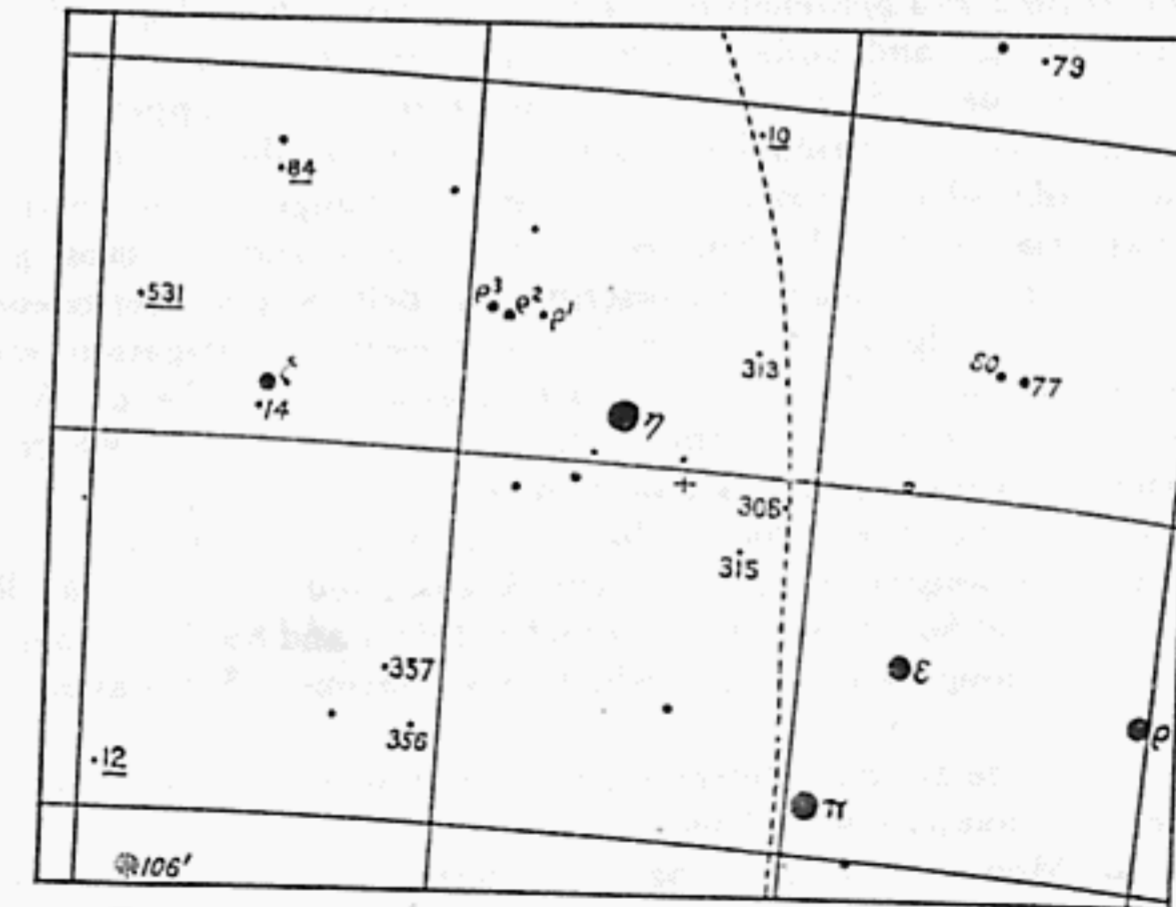


FIG. 4.—Chart of the Position of Ptolemy's Star 17 Eridani.

798. All the Greek manuscripts have the latitude erroneously $53^{\circ} 20'$; the Arabs are right. The longitudes of this and the following star are 1° too large, also by comparison with Ulugh Beg.
- 798 to 804. There is some confusion in the nomenclature of these stars, which are named v^1 to v^7 , but in different order. The designations given by Peters are those of Taylor's Madras Catalogue, the maps of the S. D. U. K. and Proctor's Atlas. The reverse order is adopted in the Uranometria Argentina, Cape Catalogues, by Peirce, Houzeau, Schjellerup, and Manitius.
802. The position agrees better with III 202, though Ulugh Beg observed III 189.
803. Ptolemy's position agrees better with Lac. *g*, though Ulugh Beg observed Lac. *f*.
804. Ptolemy's latitude is 2° or 3° too small. Sûfi's description of Ptolemy's 31-33 identifies them as Lacaille *g*, *f*, and *h*.
805. Several Greek authorities give the longitude $7^{\circ} 30'$; one reading of Paris 2389 and all the Arabs have $0^{\circ} 10'$. Halma gives as a variant $27^{\circ} 30'$, which he has clearly taken from Halley. In Paris 2394 the degrees of longitude are represented by an old cursive form of the letter ξ and so the longitude is $60^{\circ} 40'$; Grynæus has the longitude $60^{\circ} 40'$, precisely the same. Baily gives Grynæus erroneously as $7^{\circ} 40'$. All authorities, even Sûfi, designate the star as of the first magnitude. The nearest star of the first magnitude is α Eridani, which could not have been seen by Ptolemy and Sûfi. The position is near the place of θ Eridani. Peters suggests that Ptolemy's place may be a compilation from inaccurate sources; he remarks that Sûfi clearly considered θ , and not α Eridani. The computed positions of the two stars for A. D. 100 are:

	Long.	Lat.
θ Eridani...	$356^{\circ} 47'$	$-53^{\circ} 50'$
α Eridani...	$318^{\circ} 27'$	$-59^{\circ} 16'$

It is surmised that there is a large error in Ptolemy's position or that the magnitude has changed. Peters, Baily, Peirce, and Manitius identify the star as θ , Halma and Schjellerup as α Eridani; Delambre adopts Halley's longitude, $27^{\circ} 30'$, and adds in a note "*La dernière brillante du Fleuve ne peut être que la dernière de l'eau du Verseau, qui s'appelle aussi le Fleuve ou le Nil.*" θ Eridani shows no signs of variability; it is therefore highly improbable that its magnitude has changed from a first to a third magnitude star. All Almagests give mag. 1, and it is most probable that in a very ancient manuscript the $\delta = 4$ was erroneously taken to be an $\alpha = 1$, of which the present investigation shows numerous examples. Thus Ptolemy's magnitude should be 4. A corresponding error is found in the Bodleian Greek Almagest, where the magnitude of Sirius is given as 4 instead of 1.

806. Long. Baily adopts $19^{\circ} 0'$, but the authority for $19^{\circ} 40'$ is much stronger.
813. All Greek authorities give longitude $24^{\circ} 50'$; the Arabs (Bod. 369, B. M. Reg. 16, Laur. 156, and Sûfi) have $24^{\circ} 20'$, which is better and has been adopted.
822. All the manuscripts have longitude $25^{\circ} 20'$, which is erroneous. Sûfi has $20^{\circ} 20'$, which is right.
833. All authorities give longitude $23^{\circ} 0'$. Peters suggests that it should be $21^{\circ} 0'$ and it would then compare with Ulugh Beg.
836. Peters identifies as 22 Monocerotis (4.1 mag.) in preference to 19 Monocerotis (4.9 mag.); adopted by Baily and followed by Manitius, though the position of the former is more largely in error than the latter.

837. All authorities give longitude $10^{\circ} 0'$, which is 3° too large. Ulugh Beg is right. Peters has adopted $7^{\circ} 0'$.
843. Lat. The Greek manuscripts have $59^{\circ} 50'$, and the Arabs $59^{\circ} 30'$, which is better.
848. Long. The Arabs have $29^{\circ} 10'$, which is better than $29^{\circ} 30'$, as in the Greek and Baily.
849. The nomenclature of the stars in Argo is very confusing. The Index in Harvard Annals, vol. 50, has been followed as far as possible.
855. Latitude is variously given as $49^{\circ} 15'$, $49^{\circ} 30'$, $49^{\circ} 45'$, and $49^{\circ} 50'$; $49^{\circ} 30'$ seems to have the most authority. Baily adopts $49^{\circ} 15'$.
856. Latitude $49^{\circ} 30'$ of the Arabs is preferable to $49^{\circ} 50'$ of the Greeks, which Baily takes.
859. Sûfi's description leads upon Lacaille 2834. Mag. 5.3, U. A., the computed position of which is longitude $96^{\circ} 7'$, latitude $-52^{\circ} 6'$, giving errors of longitude $+127'$ and latitude $+54'$.
861. Lat. Baily gives $56^{\circ} 30'$, for which the only authority found is the Crawford Codex.
865. Lat. Greek $58^{\circ} 40'$, Arabs $58^{\circ} 20'$, the latter adopted; but this is not Ptolemy's star, whose position accords better with the group VII 102, 108, and 113.
867. Peters remarks that there is no star in the position described by Sûfi.
868. Long. Baily adopts $23^{\circ} 10'$.
869. Lat. Greek $57^{\circ} 40'$, and Arabs $57^{\circ} 0'$.
870. Peters identifies this star as Lacaille 3580, mag. 5.8, but questions whether it is not too small. There is no star in the place described by Sûfi.
875. Lat. Baily adopts $51^{\circ} 40'$.
879. Long. $14^{\circ} 10'$ has much better authority than $15^{\circ} 10'$ given by Baily.
880. Lat. All authorities agree, but it is 1° too far south.
882. Long. This is 2° too small, also by comparison with Ulugh Beg. Sûfi's description leads to Lac. 3022, which does not agree at all; longitude $113^{\circ} 2'$, latitude $-65^{\circ} 24'$.
884. Ptolemy's longitude wrong. There is no other star here larger than mag. 4.
885. The identification of this star is probably correct, but longitude is 3° in error.
886. The identification right, longitude too small.
887. Identified as *f* Carinæ, with which the position agrees, but the magnitude is 4.6, which is entirely discordant with Ptolemy's mag. 2. Baily adopts ι Argûs, but this involves an error of 12° in longitude and 3° in latitude. Schjellerup also adopts ι Argûs, the magnitude of which is 2.2 (H. R.). Is *f* Carinæ variable?
- Sûfi's description of the latter half of the constellation Argo is accurate and agrees with the sky (except Nos. 19 and 22, where there are no stars to be seen now). But the positions of Ptolemy and of Ulugh Beg do not agree with Sûfi in many places.
889. Lat. Baily has $65^{\circ} 15'$, for which there is far less authority than $62^{\circ} 15'$.
895. Lat. Baily has $13^{\circ} 40'$.
897. Lat. Paris 2389 confirms the Arabs' $14^{\circ} 45'$, which agrees better than $14^{\circ} 15'$.
898. The latitude $12^{\circ} 0'$ of the Arabs agrees better than $12^{\circ} 15'$ of the Greek.
899. Sûfi has latitude $14^{\circ} 40'$, an error in the degrees of Δ for Λ .
900. Sûfi has the erroneous latitude of $19^{\circ} 20'$.
904. The identification of this star as Ll. 18657 = W. B. 9^b 439 agrees better than Baily's star Fl. 28 A. Manitius gives it as Fl. 29.
905. All authorities have latitude $20^{\circ} 30'$, which should be $23^{\circ} 0'$. Probably it was $20^{\circ} 20'$, with the common mistake of $\text{K}\Gamma'$ for $\text{K}\Gamma$.
908. All Greeks have latitude $26^{\circ} 15'$, which is erroneous. The Arabs have it correctly, $23^{\circ} 15'$. Baily adopts $23^{\circ} 35'$ from Liechtenstein, which is an obvious mistake of Gerard of Cremona.

909. Lat. The correct latitude of $24^{\circ} 40'$ is found in the Greek manuscripts Paris 2389, 2390, Ven. 312, Vat. Reg. 90, and the Arabs. All the others, including variants in Paris 2389, 2390, and Ven. 312, have $45^{\circ} 30'$, or $49^{\circ} 30'$ (Θ for ϵ). It is possible that in a very ancient manuscript the latitude of a star in Argo was copied inadvertently into Hydra.
910. Ptolemy's longitude is 1° too small, also in comparison with Ulugh Beg.
914. Longitude 1° too large, also by comparison with Ulugh Beg.
918. Ptolemy's longitude and latitude quite erroneous. B. M. Sloane 2795 gives latitude $13^{\circ} 40'$, but probably copied from the previous star. Ulugh Beg is right.
920. All authorities give latitude $16^{\circ} 0'$, probably an error of $1s = 16^{\circ} 0'$ for $1s' = 10^{\circ} 10'$, which is adopted. Ulugh Beg's errors are similar. The position accords best with 24 Sextantis, longitude $131^{\circ} 36'$, latitude $-10^{\circ} 18'$, but that star is only mag. 6.7 (U. A.). Sûfi certainly describes 15 α Sextantis (mag. 4.5), and this star is adopted by Schjellerup and Peirce, but it assumes an error of 3° in the longitude. To all appearance there was here a star seen by Ptolemy, Sûfi, and Ulugh Beg that now is not visible or shining prominently. Manitius identifies as δ Sextantis.
927. The longitude of the Arabs has been adopted. Baily gives $1^{\circ} 20'$.
940. The large proper motion of θ Centauri, amounting in 1700 years to $28'$ in latitude, increases the discordance with Ptolemy's latitude, which is 1° too far south.
956. Peters identifies as Lac. 5390 f as Baily; Schjellerup as ξ . Sûfi calls the star double, which clearly refers to ξ^1 and ξ^2 , but the position of ξ^1 (longitude $190^{\circ} 28'$, latitude $38^{\circ} 42'$) deviates more than 5390 f .
- 962 to 971. There are very large errors in the longitude and latitude of these stars common to all the manuscripts. Some of the errors may be accidental, or due to the scribe, but the general inference is that the observations were made by different observers. (See note to 494-496.)
964. Sûfi finds no star visible near Ptolemy's place. It should be, as Sûfi remarks, of mag. 3, following upon the 29th star (No. 963). The nearest star would be Lacaille 5632 Q , but the magnitude is 5.4.
969. Long. Peters considered that there was here the not uncommon error in the Arabic of 8 for 3, which would make the longitude $213^{\circ} 20'$, but the resulting error is equally large, though of a different sign.
971. Cod. Vienna 14 and Cod. Vat. Reg. 90 give the longitude as $11^{\circ} 40'$; all other Greek sources, as well as the Arabs, give $14^{\circ} 40'$, an error of A for Δ . The adoption of $11^{\circ} 40'$ would give a more consistent error in Ptolemy's longitude = $+2^{\circ} 43'$.
- 979 to 981. The errors in longitudes of these three stars differ from all others in the constellation Lupus in that they have a *minus* sign. From this Peters inferred that they may have been derived from a different observer.
982. Long. The Greeks $22^{\circ} 0'$ and the Arabs $20^{\circ} 20'$. Peters corrects it to $26^{\circ} 0'$. The identification of this star presents considerable difficulty. The description states "Australis de tribus quæ sunt in extrema cauda." The following star, 983, correctly identified as ι Lupi, is "Media ipsarum," and the next, 984, also correctly identified as τ^1 and τ^2 , is "Borealis ipsarum." Peters first suggested that the star was Lac. 5209, but this is in Crux and a long way from the described position. Sûfi could not find the star and of course it is omitted by Ulugh Beg. Peters finally adopted Lac. 6003 ρ , which, assuming an error of 4° in longitude, would agree well; but the position is quite discordant with the description. Manitius identifies 982-984 as σ , ρ , and α Lupi, the positions of which would accord with the description, but involve very large errors in longitude; moreover, α Lupi

- seems well identified as Ptolemy's second star in Lupus. Baily's identification for the three stars is Lac. 1201 τ , 1215 ι , and 1209 κ (1201 = ι and 1209 = τ^1). It must remain a question whether there is here a variable or a lost star.
983. Long. Baily has $24^{\circ} 50'$.
989. Peters' identification agrees with Baily and Manitius. The longitude and latitude of the Arabs has been adopted. Baily gives longitude $27^{\circ} 10'$, latitude $11^{\circ} 50'$.
990. All authorities give longitude $26^{\circ} 30'$, except B. M. Arabic 7475, which has $27^{\circ} 30'$. Halma has $27^{\circ} 30'$, which would be much better. Peters questions his authority. There is no doubt that Halma took it from Halley's edition. Ulugh Beg's longitude is also 1° too small. The latitudes of the last three stars in Lupus are 1° too far north.
992. Several Greek authorities have longitude $3^{\circ} 0'$; the Arabs $0^{\circ} 20'$; the former is adopted—an error of $\Gamma' = 0^{\circ} 20'$ for $\Gamma = 3^{\circ} 0'$. Baily adopts $3^{\circ} 10'$.
993. Long. Baily adopts $26^{\circ} 20'$.
994. Lat. With the exception of Ven. 311, Laur. 1, and Laur. 6, all Greek codices, as well as Grynæus and Halma, have latitude $1^{\circ} 30'$ instead of $30^{\circ} 30'$. An error of $A = 1^{\circ}$ for $\Lambda = 30^{\circ}$.
997. Latitude $34^{\circ} 0'$ adopted from the Arabs. Baily has $34^{\circ} 15'$. Peters agrees with Schjellerup in the identification of the stars in Ara. Baily identifies in this order: γ , ϵ , δ , α , β , η , θ . There is a large error in all the longitudes, averaging $2^{\circ} 18'$ too small. These errors resemble those already referred to under 494-496, and 962-971.
998. Peters identifies as $\frac{\delta^1 + \delta^2}{2}$ Telescopii, as it agrees better in longitude, but remarks that it is not probable that α Telescopii should have been omitted.
1000. Baily identifies as 1566 ζ , which star Peters identifies in No. 1001.
1001. Baily identifies as β . Peters considers β to be 1003.
1004. Longitude $16^{\circ} 50'$ adopted from Cod. Vat. 1594, and the Arabs. Baily has $16^{\circ} 20'$.
1008. Baily has latitude $15^{\circ} 50'$, for which there is no authority; it is probably a misprint.
1009. Identified as Lac. 7748, which agrees better with Ulugh Beg's observations than Lac. 7758 = 1528 κ , identified by Baily, Schjellerup, and Manitius. Sûfi's description refers clearly to Lac. 7748.
1011. This is the same star as No. 670. Baily gives latitude $23^{\circ} 0'$, though for No. 670 he has $20^{\circ} 20'$.
- 1013 and 1015. Vatican Reg. 90 gives the longitudes as $30^{\circ} 10'$ and $30^{\circ} 20'$, respectively. Probably the original degrees were $\Delta = 4$, then erroneously $A = 1$, then erroneously $\Lambda = 30$.
1017. Peters remarks that longitude $2^{\circ} 10'$, adopted by Halma, would be much better, but there is no authority. Here again Halma has taken the longitude from Halley, which, as already pointed out, is not a collation of any manuscripts, but an edition in which many errors are corrected by computation.
1023. Baily has taken the Greek description of this star from Grynæus, which is identical with Paris 2389; both are erroneous, as they omit the word $\tau\rho\iota\omega\nu$. Vatican 1594 is correct.
- 1023 to 1028. Peters identifies these six stars, forming the *informata* of Piscis Austrinus, as Lacaille 8579, 8639, 8761, 8685, 8731, and 8689. The identifications are not open to much doubt, but there are large errors in the coördinates of the six stars, averaging in longitude $+1^{\circ} 21'$, and in latitude $+6^{\circ} 36'$. Upon this identification 1028 is the same star as 613.

TABLE VI.

Differences of Identification.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
URSA MAJOR.						
18	10	30 φ	44 Lyncis	φ	63 Heis.	φ
40	Inf. 5	10 Leo minor	?10 Leo minor	10 Leo minor		11 Leo minor.
41	Inf. 6	IX 115	—	—		8 Leo minor.
42	Inf. 7	36 Lyncis	—	—		10 Leo minor.
CEPHEUS.						
86	Inf. 1	μ Cephei	?XXI 248	ν Cephei	μ	μ
BOOTES.						
97	10	2 η Cor. Bor.	?1 or 2 Cor. Bor.	η Cor. Bor.	η Cor. Bor.	χ
98	11	1 ο Cor. Bor.	48 χ	χ		—
99	12	41 ω	45 c	b		c
100	13	46 b	43 ψ	ω		ψ
101	14	43 ψ	46 b	ψ		b
102	15	45 c	41 ω	c		ω
HERCULES.						
138	20	74	77 x	x		x
139	21	77 x	82 y	y		y
140	22	82 y	88 z	z		z
LYRA.						
156	8	9 ν ²	8 ν ¹	ν	8	ν
CASSIOPEIA.						
184	7	35 (Hev.) ε	?II 72	ε		ε
185	8	33 θ	33 θ	μ		μ
186	9	34 φ	34 φ	θ		θ
PERSEUS.						
196	6	18 (Hev.) ε	II 253 ε	ε		ε
218	Inf. 2	14 (Hev.) Camel.	IV 7	12 Hev. Camel.		34 Hev. Camel.
AURIGA.						
227	8	10 η	10 η	ξ		η
228	9	8 ξ	8 ξ	η		ξ
233	14	14	4	η	4	ξ
OPHIUCHUS.						
246	13	40 ξ	—	40	40	ξ
247	14	36 λ	—	36	—	λ
248	15	42 θ	—	θ	θ	θ
249	16	44 b	—	44	7 Behr.	b
250	17	51 c	—	51	—	51
251	18	{52} Sagittarii	52	58	—	58
AQUILA.						
289	4	59 ξ	54 ο	ξ	ο	ο
290	5	50 γ	50 γ	ν	γ	γ

PTOLEMY'S CATALOGUE OF STARS.
Differences of Identification—continued.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
DELPHINUS.						
308	8	3 η	3 η	ξ		η
309	9	4 ζ	4 ζ	η		ζ
EQUULEUS.						
311	1	8 α	8 α	—		α
312	2	10 β	10 β	—		β
313	3	5 γ	5 γ	—		γ
314	4	7 δ	7 δ	—		δ
PEGASUS.						
327	13	50 ρ	50 ρ	σ		ρ
328	14	49 σ	49 σ	ρ		σ
ANDROMEDA.						
355	21	49 λ	49 λ	λ		ξ
356	22	52 χ	52 χ	χ		ω
357	23	1 ο	2	ο	ο	ο
TAURUS.						
401	22	69 υ ¹	69 υ ¹	υ		υ ²
403	24	37 λ ¹	—	λ		λ
404	25	50 ω ²	—	ω	43	ω
409	30	19 ε	19	—	ω	16
410	31	23 δ	23	—	—	17
411	32	{25 η 27 f}	{27	—	—	η
412	33	III 170	18	—	—	—
415	Inf. 3	109 n	105	n	—	38 H.
417	Inf. 5	126	126	—	—	n
418	Inf. 6	129	128	119	—	130
419	Inf. 7	121	121	Σ730	129	129
GEMINI.						
432	9	58	76 c	b	52 Tauri	63
445	Inf. 4	85	85	85	—	g
446	Inf. 5	81 g	81 g	g	—	f
447	Inf. 6	74 f	74 f	f	—	68
448	Inf. 7	16 ζ Cancri	16 ζ Canc.	—	—	ζ Canc.
CANCER.						
458	Inf. 1	{62 ο ¹ 63 ο ²	{81 π ¹	π	π ¹	π
460	Inf. 3	69 ν	69 ν	ξ	—	ν
461	Inf. 4	77 ξ	77 ξ	ν	—	ξ
LEO.						
482	21	—	71	72	—	θ
483	22	70 θ	70 θ	θ	—	θ
486	25	84 τ	84 τ	τ	—	τ
494	Inf. 6	15 c Comæ	—	p ⁵	—	τ
495	Inf. 7	7 h Comæ	4 Comæ	15 Comæ	—	15 Comæ.
496	Inf. 8	23 k Comæ	21 Comæ	12 Comæ	—	7 Comæ.
VIRGO.						
504	8	46	46	k	46	k
512	16	74 l	74 l ²	l	—	l ²
514	18	82 m	82 m	LL. 25396	82 m	m
515	19	68 i	68 i	LL. 25086	—	i
528	Inf. 6	89	73	89	89	89

PTOLEMY'S CATALOGUE OF STARS.
Differences of Identification—continued.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
LIBRA.						
541	Inf. 5	43 κ	41	η		41
542	Inf. 6	O. Arg. 14782	43 κ	κ		κ
544	Inf. 8	39	39	3 Hev. Scorp.		2 H. Scorp.
545	Inf. 9	40	40	o Scorp.		o Scorp.
SCORPIUS.						
560	15	XVI 206 ζ ²	ζ ²	—		ζ ¹
567	Inf. 1	γ Telescopii	—	—	65 Behr.	G
569	Inf. 3	3 Sagittarii	44 Oph. or 3 Sag.	3 Sagittarii	44 Oph.	43 Oph.
SAGITTARIUS.						
586	17	56 f	56 f	f		57
594	25	XVIII 17	η	β Telescopii		η
595	26	XIX {330 k ¹ 333 k ²	θ	m Lac.		θ ¹
596	27	XIX 297 t	t	e Lac.		t
CAPRICORNUS.						
615	15	36 b	35	b		b
AQUARIUS.						
634	6	13 v	13 v	Fl. 7		v
645	17	38 e	37 e ¹	e		e
649	21	68 g ²	59 v	68		g ²
651	23	63 κ ?	67 ?	67		κ
655	27	92 χ	92 χ	—		χ
656	28	91 ψ ¹	91 ψ ¹	—		ψ ¹
658	30	94	94 ?	97		131 H
659	31	102 ω ¹	102 ω ¹	—		ω ¹
662	34	106 i ¹	104 A ²	i ¹		i ¹
663	35	108 i ²	106 i ¹	i ²		i ²
666	38	101 b ¹	101 b ¹	b ²		b ²
PISCES.						
704	31	81 ψ ³	81 ψ ³	ψ ³		χ
707	34	84 χ	84 χ	χ		99 H
CETUS.						
716	5	—	78 v	ξ ²		v
717	6	—	73 ξ ²	μ		μ
728	17	19 φ ²	21	19	21	φ ²
729	18	O. 198	19 φ ²	23		φ ²
730	19	17 φ ¹	—	17	28 Heis Ceti.	φ ¹
731	20	O. 161	17 φ ¹	18		φ ¹
ORION.						
742	9	72 f ²	72 f ²	f ¹		f ²
743	10	69 f ¹	69 f ¹	f ²		f ¹
744	11	54 χ ¹	57 χ ²	χ ¹		χ ¹
745	12	62 χ ²	64 χ ³	χ ²		χ ²
748	15	33 π ¹	33 π ¹	ψ		π ¹
749	16	30 ψ ²	30 ψ ²	25		ψ
752	19	9 o ²	6 g	o ²		o ²
753	20	7 π ¹	7 π ¹	π ¹		π ¹
755	23	1 π ³	1 π ³	π ⁴		π ³
756	24	3 π ⁴	3 π ⁴	π ⁵		π ⁴
763	30	{42} c {45} c	42 c	v		c

PTOLEMY'S CATALOGUE OF STARS.
Differences of Identification—continued.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
ERIDANUS.						
779	8	40 o ²	40 d	o ²	98 Heis.	o ²
787	16	3 η	3 η	η		η
788	17	—	σ	LL. 4969		ρ ³
798	27	50 v ^{6,7,1}	50 v ⁶	v ¹		η
800	29	43 v ^{6,13}	43 v ⁶	v ¹		v ¹
802	31	III 202 v ²	v ²	v ²		d
803	32	III 189 v ²	v ²	v ²	58 Behr.	v ⁶
804	33	III 149 v ¹	v ¹	h		g
805	34	{II 238} θ {III 239} θ	θ	a		h
CANIS MAJOR.						
825	8	7 v ²	6 v ¹	v ²		v ²
836	Inf. 1	22 Monocerotis	19 Monoc.	22 Monoc.		19 Monoc.
837	Inf. 2	VI 9 θ Columb.	485 Lac.	θ Columb.		θ Columb.
838	Inf. 3	VI 65 κ Columb.	497 Lac.	κ Columb.		κ Columb.
839	Inf. 4	VI 95 δ Columb.	510 Lac.	δ Columb.		δ Columb.
840	Inf. 5	VI 136 λ Can. maj.	521 Lac.	λ Can. maj.		λ Can. maj.
841	Inf. 6	V 238 μ Columb.	444 Lac.	μ Columb.		μ Columb.
842	Inf. 7	V 276 λ Columb.	453 Lac.	—		λ Columb.
843	Inf. 8	V 297 γ Columb.	465 Lac.	γ Columb.		γ Columb.
844	Inf. 9	V 267 β Columb.	452 Lac.	β Columb.		β Columb.
845	Inf. 10	V 196 α Columb.	434 Lac.	α Columb.		α Columb.
846	Inf. 11	V 140 ε Columb.	419 Lac.	ε Columb.		ε Columb.
ARGO NAVIS.						
857	9	VII 200 i Pup.	σ	σ		17 H. Arg.
859	11	{VII 99 VII 108}	v	VII 137		3 H. Arg.
860	12	VII 68 π Pup.	λ	π		π Pup.
861	13	VII 172 f Pup.	f	—	f Pup.	f Pup.
862	14	VII 186 {d ¹ d ² } Pup.	φ ¹	—	d Pup.	d ¹ Pup.
863	15	VII 214 c Pup.	φ ²	—	c Pup.	c Pup.
864	16	VII 254 b Pup.	ψ ? b	—	b Pup.	b Pup.
865	17	VII 306 ζ Pup.	δ	—	ζ Pup.	ζ Pup.
866	18	VII 253 α Pup.	ω ¹	α		—
867	19	Lac. 3128	ω ²	—		—
868	20	VIII 21 h ¹ Pup.	A ¹	r		h ¹ Pup.
869	21	VIII 35 h ² Pup.	A ²	q		h ² Pup.
870	22	Lac. 3580	p ¹	—	p ¹ 52 Behr.	d Vel.
871	23	VIII 168 d Vel.	p ²	—	p ² d Vel.	a Vel.
872	24	VIII 139 e Vel.	p ³	—	p ³ e Vel.	b Vel.
873	25	VIII 176 a Vel.	a	—	a Vel.	D Vel.
874	26	VIII 155 b Vel.	b	—	b Vel.	C Vel.
875	27	VIII 145 o ¹ (β Pyx.) (b Mali.)	o ¹	—	b Mali.	β Pyx.
879	31	IX 1 λ Vel.	ε	λ		λ Vel.
881	33	VII 135 σ Pup.	i	σ		—
882	34	VII 235 P Pup.	r	P		—
883	35	γ Vel.	ζ	γ		γ Vel.
884	36	χ Car.	η	χ		ε Car.
885	37	o Pup.	q	δ	δ Arg.	δ Vel.
886	38	δ Vel.	θ	κ		κ Vel.
887	39	f Car.	v	t		φ Vel.

PTOLEMY'S CATALOGUE OF STARS.
Differences of Identification—continued.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
ARGO NAVIS—cont.						
888	40	κ Vel.	b	N	b κ.	ν Car.
889	41	N. Vel.	c	φ	φ Arg.	θ Car.
890	42	V 315 η Columb.	471 Lac.	η Columb.	τ Pup.	τ Pup.
891	43	VI 205 ν Pup.	g	ν Arg.	g ν.	σ Pup.
893	45	τ Pup.	h	τ Pup.	h τ.	τ Pup.
HYDRA.						
904	11	{LL. 18657 W 9 ^h 439}	28 A	—	—	29
906	13	38 κ.	38 κ.	ν ¹	—	κ
907	14	39 ν ¹	39 ν ¹	ν ²	—	ν ¹
908	15	40 ν ²	40 ν ²	λ.	—	λ ²
919	Inf. 1	30 Monocerotis.	1	30 Monoc.	—	30 Monoc.
920	Inf. 2	{24 Sextantis. 15 a Sextantis.}	—	15 Sextantis.	15 Sextantis.	δ Sextantis.
CRATER.						
924	4	27 ζ.	27 ζ.	η.	ζ.	ζ
925	5	14 ε.	14 ε.	θ.	—	ε
926	6	30 η.	30 η.	ζ.	η.	η
927	7	21 θ.	21 θ.	ε.	—	θ
CENTAURUS.						
941	7	XIII 99 d.	1150 Lac. ψ.	d	—	d
942	8	XIV 40 ψ.	1205 Lac. l.	ψ	—	ψ
943	9	XIV 55 a.	1207 Lac. o.	a	—	a
944	10	XIV 150 c ¹ .	1234 Lac. π.	c	—	c ¹
945	11	XIV 141 b.	1232 Lac. ρ.	b	—	c ²
946	12	XIII 197 ν.	1165 Lac. τ.	ν	τ (ν).	ν
947	13	XIII 198 μ.	1166 Lac. v.	μ	ν (μ).	μ
948	14	XIII 246 φ.	1182 Lac. φ.	φ	φ (φ).	φ
949	15	XIII 288 χ.	1191 Lac. m.	χ	m (χ).	χ
950	16	XIV 109 η.	1219 Lac. κ.	η	—	η
951	17	XIV 216 κ.	1255 Lac. σ.	κ	—	κ
952	18	XIII 231 ζ.	1177 Lac. λ.	ζ	λ (ζ).	ζ
953	19	XIII 267 ν ² .	1184 Lac. η.	o.	—	ν ¹
954	20	XIII 249 ν ¹ .	1183 Lac. χ.	—	—	ν ²
955	21	ω.	1148 Lac. ω.	—	—	ω
956	22	f.	1123 Lac. o.	ξ.	ξ.	f
957	23	γ.	1098 Lac. μ.	γ.	—	γ
958	24	τ.	1093 Lac. c.	τ.	—	ρ
959	25	σ.	1086 Lac. p.	σ.	—	δ
960	26	δ.	1064 Lac. β.	δ.	—	—
961	27	ρ.	1068 Lac. e.	ρ.	—	—
962	28	M.	1155 Lac. δ.	—	—	—
963	29	ε.	—	ε.	—	ε
964	30	Q.	—	—	—	γ Crucis.
965	31	γ Crucis.	1070 Lac. ν.	γ Crucis.	—	δ Crucis.
966	32	β Crucis.	1107 Lac. ξ.	β Crucis.	—	β Crucis.
967	33	δ Crucis.	1025 Lac. f.	δ Crucis.	δ Crucis.	α Crucis.
968	34	α Crucis.	1082 Lac. ζ.	α Crucis.	α Crucis.	λ
970	36	β Cent.	1185 Lac. γ.	β Cent.	—	β Cent.
971	37	μ Crucis.	1107 Lac. ε.	θ Cent.	θ.	—

PTOLEMY'S CATALOGUE OF STARS.
Differences of Identification—continued.

Baily's No.	Ptolemy's No.	Peters.	Baily.	Schjellerup.	Peirce.	Manitius.
LUPUS.						
972	1	XIV 211 β.	1254 Lac. o.	β.	—	β
973	2	a.	1231 Lac. a.	a.	—	o
974	3	XV 31 δ.	1283 Lac. ζ.	δ.	—	γ
975	4	XV 98 γ.	1293 Lac. η.	γ.	—	δ
976	5	XV 35 ε.	1285 Lac. θ.	ε.	—	ε
977	6	λ.	1263 Lac. π.	λ.	—	π
978	7	XV 242 π.	1258 Lac. β.	π.	—	κ
979	8	μ.	1274 Lac. ξ.	μ.	—	μ
980	9	κ.	1266 Lac. ρ.	κ.	—	ν
981	10	ζ.	1265 Lac. σ.	ζ.	—	ζ
982	11	ρ?	1201 Lac. τ.	—	—	σ
983	12	ι.	1215 Lac. ι.	ι.	—	ρ
984	13	XIV {66τ ¹ . 67τ ² }. 1209 Lac. κ.	—	τ.	κ (τ).	a
985	14	XV 217 η.	1325 Lac. ν.	η.	—	η
986	15	XV 248 θ.	1335 Lac. μ.	θ.	μ (ξ).	θ
987	16	XV 174 Fl. 5 χ.	3 γ.	λ.	—	ψ.
988	17	XV 204 ξ.	5 λ.	—	—	χ
989	18	XV 10 Fl. 1 i.	1 ε.	δ.	ε 30 Behr.	i
990	19	XV 22 Fl. 2 f.	2 δ.	—	δ 33 Behr.	f
ARA.						
991	1	XVII 125 σ.	γ.	σ.	—	σ
992	2	θ.	ε.	θ.	—	θ
993	3	a.	δ.	a.	δ (a).	a
994	4	ε ¹ .	α.	ε.	—	ε
995	5	γ.	β.	γ.	a (ε ¹).	γ
996	6	β.	η.	β.	—	β
997	7	ζ.	θ.	ζ.	—	ζ
CORONA AUSTRALIS.						
998	1	XVIII {73 δ ¹ . 76 δ ² } Tel.	a.	θ.	—	a Teles.
999	2	XVIII {166 η ¹ . 169 η ² }.	ε.	η.	—	η
1000	3	Lac. 7909.	ζ.	o.	—	—
1001	4	XVIII 250 ζ.	β.	ζ.	—	ζ
1002	5	XVIII 291 δ.	η.	δ.	—	δ
1003	6	XVIII 305 β.	θ.	β.	—	β
1004	7	XVIII 300 a.	γ.	κ.	γ.	a
1005	8	XVIII 280 γ.	δ.	γ.	—	γ
1006	9	XVIII 230 ε.	μ.	—	—	ε
1007	10	XVIII 222 ν.	ν.	—	—	—
1008	11	XVIII 142 λ.	ι.	λ.	—	λ
1009	12	Lac. 7748 ξ Bode.	κ.	κ Bode.	—	κ
1010	13	XVIII 85 θ.	λ.	ξ Bode.	—	θ
PISCIS AUSTRINUS.						
1022	Inf. 12	XXI 308 γ Gruis.	κ.	γ.	—	γ Gruis.
1023	Inf. 1	XX 307 α Mic.	1694 Lac.	λ Gruis.	—	α Mic.
1024	Inf. 2	XX 403 γ Mic.	1717 Lac.	μ Gruis.	—	γ Mic.
1025	Inf. 3	XXI 46 ε Mic.	—	δ Gruis.	—	ε Mic.
1026	Inf. 4	XX 445.	1704 Lac.	β Gruis.	—	δ Mic.
1027	Inf. 5	XXI 12.	—	α Gruis.	—	—
1028	Inf. 6	24 A Capric.	4.	ι Gruis.	4.	—

THE STAR MAGNITUDES.

The magnitudes of the stars in the catalogue are those deduced as most probable from consideration of the Table of Star Magnitudes (pp. 122-143), besides many other authorities mentioned in the notes.

The magnitudes in the Greek codices generally agree very well. Comparing the two oldest Greek codices, Paris 2389 and Vatican 1594, twelve differences are found, of which Paris 2389 is correct in ten and Vatican 1594 in two cases. Comparing Vatican 1594 with Venice 313, only 4 differences are noted. Comparing Paris 2389 with the Arabic codex, British Museum Reg. 16, there are 35 differences, of which Paris 2389 is correct in 21 and B. M. Reg. 16 correct in 13 cases, with one case in which both are probably wrong. The Arabic codex, B. M. Reg. 16, is particularly valuable from the great care with which it has been written. In all series of stars of the same magnitude, the magnitudes of the first and last only are written—a method which avoids many mistakes.

The magnitudes adopted in the catalogue differ from those in Paris 2389 in the following 14 stars: Baily, Nos. 128, 129, 130, 154, 211, 352, 480, 509, 576, 736, 764, 765, 824, and 885.

It will be seen in Table VIII that Dr. Peters has adopted magnitudes for some stars which differ from all manuscripts of the Almagest yet examined, and for which no authority can be found. In a note on one of his collations, he says that he has "inserted the *revised* magnitudes of the Paris Codex 2389, besides several notes on the stars in my copy of Baily's Ptolemy" (Mems. R. A. S., Vol. XIII), but unfortunately this volume can not be found.

The magnitudes in Ptolemy's catalogue have been fully discussed by Prof. E. C. Pickering in H. A., Vol. XIV, Part II. In this memoir he has reduced Ptolemy's magnitudes to the photometric scale of the Harvard Photometry, and arrives at the accompanying photometric values:

Ptolemy magnitude.	Photometric magnitude.	Ptolemy magnitude.	Photometric magnitude.
1	0.5	3-4	3.8
1-2	1.2	4-3	3.8
2-1	1.2	4	4.4
2	2.1	4-5	4.6
2-3	2.6	5-4	4.7
3-2	2.7	5	5.0
3	3.3	6	5.4

In the following table of whole magnitudes 2 to 6 (Table VII), a rather larger number of stars is employed and the magnitudes are based on the Harvard Revised Photometry. The corresponding figures from H. A., Vol. XIV, are appended in italics. It will be seen that the results do not suggest any material difference from those obtained by Professor Pickering in the above investigation.

TABLE VII.

Ptolemy magnitude.	No. of stars.				Mean magnitudes.			
	North.	Zodiac.	South.	All.	North.	Zodiac.	South.	All.
2	11	6	12	29	2.20	2.10	2.14	2.14
	<i>12</i>	<i>6</i>	<i>7</i>	<i>25</i>	<i>2.04</i>	<i>1.95</i>	<i>2.23</i>	<i>2.07</i>
3	63	52	46	161	3.22	3.24	3.35	3.27
	<i>58</i>	<i>44</i>	<i>34</i>	<i>136</i>	<i>3.28</i>	<i>3.31</i>	<i>3.36</i>	<i>3.31</i>
4	121	100	111	332	4.32	4.45	4.30	4.36
	<i>119</i>	<i>105</i>	<i>75</i>	<i>299</i>	<i>4.33</i>	<i>4.48</i>	<i>4.32</i>	<i>4.38</i>
5	48	95	38	181	4.84	5.08	4.64	4.85
	<i>40</i>	<i>82</i>	<i>16</i>	<i>138</i>	<i>4.81</i>	<i>5.04</i>	<i>4.82</i>	<i>4.95</i>
6	13	24	8	45	5.27	5.36	5.22	5.28
	<i>9</i>	<i>25</i>	<i>4</i>	<i>38</i>	<i>5.46</i>	<i>5.38</i>	<i>5.18</i>	<i>5.38</i>

In Table VIII the first column gives the number of the star in Baily's Ptolemy; the second column the name of the star; the third gives the magnitudes assigned by Dr. Peters, an asterisk (*) indicating those which differ from the magnitudes adopted in the catalogue; the next three columns give the magnitudes in the Greek codices, Paris 2389, Vatican 1594, and Venice 313; the following column gives the magnitudes adopted by Manitius from the several Greek manuscripts he examined; then follow the magnitudes in three Arabic codices, British Museum Reg. 16, British Museum 7475, and Bodleian 369; and in the last column is given the magnitudes in the Harvard Revised Photometry; for double stars the combined magnitude is given.

The Notes on pp. 144-150 give the variants from the adopted magnitudes, in the Greek codices, Paris 2389, Vatican 1594, Vatican 1038, Venice manuscripts 302, 310, 312, and 313, and Laurentian 48; the Latin codex Laurentian 6, and the three Arabic codices, British Museum Reg. 16 and 7475, and Bodleian 369. The magnitudes in the Latin manuscripts of Gerard of Cremona (Laurentian 45 and British Museum, Sloane 2795) show so many discordances that they are passed over, except in a few instances. Baily has omitted the qualifying words *μείζων* and *ελάσσων*, consequently the variants in his edition refer only to magnitudes not so qualified in the catalogue.

PTOLEMY'S CATALOGUE OF STARS.
Star Magnitudes—continued.

Baily's No.	Name.	Peters.	Greek.				Arabic.			Harv. R. P.
			Paris 2389.	Vatican 1594.	Venice 313.	Manitius printed.	B. M. Reg. 16.	B. M. 7475.	Bod 369.	
ARGO NAVIS—cont.										
857	VII 200 = I Pup	4	4	4	4	4	4	4	4.8	
858	VII 277	4	4	4	4	4	4	4	6.5	
859	{ VII 99 VII 108 }	4	4	4	4	4	4	4	5.0	
860	VII 68 π Pup	3	3	3	3	3	3	3	2.7	
861	VII 172 f Pup	5	5	5	5	5	5	5	4.6	
862	VII 186 { d ¹ Pup d ² Pup d ³ Pup }	5	5	5	5	5	5	5	4.2	
863	VII 214 c Pup	4-5*	4	4	4	4	4	4	3.7	
864	VII 254 b Pup	4	4	4	4	4	4	4	4.5	
865	VII 306 ζ Pup	2	2	2	2	2	2	2	2.3	
866	VII 253 a Pup	5	5	5	5	5	5	5	3.8	
867	Lac. 3128	5	5	5	5	5	5	5	5.5	
868	VIII 21 h ¹ Pup	5	5	5	5	5	5	5	4.4	
869	VIII 35 h ² Pup	5	5	5	5	5	5	5	4.4	
870	Lac. 3580	4-3	4-3	4-3	4-3	4-3	4-3	4-3	5.8	
871	VIII 168 d Vel	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.1	
872	VIII 139 e Vel	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.1	
873	VIII 176 a Vel	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.1	
874	VIII 155 b Vel	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.1	
875	VIII 145 β Pyx	3	3	3	3	3	3	3	4.0	
876	VIII 162 α Pyx	3-4*	3	3	3	3	3	3	3.7	
877	VIII 193 γ Pyx	4	4	4	4	4	4	4	4.2	
878	VIII 220 δ Pyx	4	4	4	4	4	4	4	4.9	
879	IX 1 λ Vel	2	2	2	2	2	2	2	2.2	
880	IX 116 ψ Vel	2-3	2-3	2-3	2-3	2-3	2	2-3	3.6	
881	VII 135 σ Pup	4	4	4	4	4	4	4	3.3	
882	VII 235 P Pup	6	6	6	6	6	6	6	4.2	
883	γ Vel	2	2	2	2	2	2	2	2.2	
884	χ Car	2	2	2	2	2	2	2	3.6	
885	o Pup	2*	2	2	2	3	3	3	4.6	
886	δ Vel	3	3	3	3	3	3	3	2.0	
887	f Car	2	2	2	2	2	2	2	4.6	
888	κ Vel	3-4*	3	3	3	3	3	3	2.6	
889	N Vel	3	3	3	3	3	3	3	3.0	
890	V 315 η Columb	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.0	
891	VI 205 ν Pup	3-2	3-2	3-2	3-2	3-2	3-2	3-2	3.2	
892	α Arg. Canopus	1	1	1	1	1	1	1	-0.8	
893	τ Pup	3-2	3-2	3-2	3-2	3-2	2	3-2	2.8	
HYDRA.										
894	5 σ	4	4	4	4	4	4	4	4.5	
895	4 δ	4	4	4	4	4	4	4	4.2	
896	11 ε	4	4	4	4	4	4	4	3.5	
897	7 η	4	4	4	4	4	4	4	4.3	
898	16 ζ	4	4	4	4	4	4	4	3.3	
899	18 ω	5	5	5	5	5	5	5	5.4	
900	22 θ	4	4	4	4	4	3	4	3.8	
901	32 π ²	4	4	4	4	4	4	4	4.5	
902	35 ι	4	4	4	4	4	4	4	4.1	
903	31 π ¹	4	4	4	4	4	4	4	4.8	
904	LL. 18657	6	6	6	6	6	6	6	5.4	
905	30 α	2	2	2	2	2	2	2	2.2	

PTOLEMY'S CATALOGUE OF STARS.
Star Magnitudes—continued.

Baily's No.	Name.	Peters.	Greek.				Arabic.			Harv. R. P.
			Paris 2389.	Vatican 1594.	Venice 313.	Manitius printed.	B. M. Reg. 16.	B. M. 7475.	Bod. 369.	
HYDRA—continued.										
906	38 κ	4	4	4	4	4	3	4	5.0	
907	39 υ ¹	4	4	4	4	4	4	4	4.3	
908	40 υ ²	4	4	4	4	4	4	4	4.7	
909	42 μ	3-4*	3	3	3	3	3	3	4.1	
910	φ (2 Crat.)	4	4	4	4	4	4	4	5.1	
911	ν (4 Crat.)	3	3	3	3	3	3	3	3.3	
912	(11 β Crat.)	4-3	4-3	4-3	4-3	4	4	4	4.5	
913	χ ¹ (9 Crat.)	4	4	4	4	4	4	4	5.1	
914	ξ (19 Crat.)	4	4	4	4	4	4	4	3.7	
915	ο (25 Crat.)	4	4	4	4	4	4	4	4.9	
916	β (28 Crat.)	3	3	3	3	3	3	3	4.4	
917	46 γ	4-3	4-3	4-3	4-3	4-3	4-3	4-3	3.3	
918	49 π	4-3	4-3	4-3	4-3	4-3	4-3	4-3	3.5	
919	30 Monoc	3	3	3	3	3	3	3	3.9	
920	24 Sextantis 15 α Sextantis	3	3	3	3	3	3	3	6.7 4.5	
CRATER.										
921	7 α	4	4	4	4	4	4	4	4.2	
922	15 γ	4	4	4	4	4	6	4	4.1	
923	12 δ	4-5*	4	4	4	4	4	4	3.8	
924	27 ζ	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.9	
925	14 ε	4	4	4	4	4	4	4	5.1	
926	30 η	4-5	4-5	4-5	4-5	4-3	4-3	4-3	5.2	
927	21 θ	4	4	4	4	4	4	4	4.8	
CORVUS.										
928	1 α	3	3	3	3	3	3	3	4.2	
929	2 ε	3	3	3	3	3	3	3	3.2	
930	5 ζ	5	5	5	5	5	5	5	5.3	
931	4 γ	3	3	3	3	3	3	3	2.8	
932	7 δ	3	3	3	3	3	3	3	3.1	
933	8 η	4	4	4	4	4	4	4	4.4	
934	9 β	3	3	3	3	3	3	3	2.8	
CENTAURUS.										
935	2 g	5-4	5-4	5-4	5-4	5-4	5-4	5-4	4.4	
936	4 h	5-4	5-4	5-4	5-4	5-4	5-4	5-4	4.8	
937	1 i	4-5*	4-3	4-3	4-3	4-3	5-4	4-3	4.4	
938	3 k	5-4	5-4	5-4	5-4	5-4	5-4	5-4	4.7	
939	XIII 53 l	3	3	3	3	3	3	3	2.9	
940	5 θ	3	3	3	3	3	3	3	2.3	
941	XIII 99 d	4	4	4	4	4	4	4	4.0	
942	XIV 40 ψ	4	4	4	4	4	4	4	4.2	
943	XIV 55 a	4	4	4	4	4	4	4	4.5	
944	XIV 150 c ¹	4	4	4	4	4	4	4	4.1	
945	XIV 141 b	4	4	4	4-5	4	4	4	4.1	
946	XIII 197 ν	4-3	4-3	4-3	4-3	4-3	4-3	4-3	3.5	
947	XIII 198 μ	4-3	4-3	4-3	4-3	4-3	4-3	4-3	3.3	
948	XIII 246 φ	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.0	
949	XIII 288 χ	4-3	4-3	4-3	4-3	4-3	4-3	4-3	4.5	
950	XIV 109 η	3	3	3	3	3	3	3	2.6	
951	XIV 216 κ	4	4	4-3	4	4	4	4	3.3	
952	XIII 231 ζ	3-2	3-2	3	3-2	3-2	3-2	3-2	3.1	

NOTES ON THE STAR MAGNITUDES.

Baily's No.	Star.	Notes.
7	URSA MINOR. 7. 13 γ	All manuscripts give mag. 2. Peters adopts mag. 3, found only in Sûfi.
15 18	URSA MAJOR. 7. 14 τ 10. 30 φ	All manuscripts have mag. 4. No authority for Peters' mag. 4-5. Vat. 1594, Vat. 1038, Ven. 310, Ven. 313, and Laur. 48 have mag. 4. The adopted mag. 4-5 is from Paris 2389, Ven. 302, Ven. 312, Vat. Reg. 90, Laur. 6, and all the Arabs.
20 21 28 29 30 31 32 35	12. 9 ι 13. 12 κ 20. 33 λ 21. 34 μ 22. 52 ψ 23. 54 ν 24. 53 ξ 27. 85 η	All manuscripts have mag. 3. No authority for Peters' mag. 3-4. All manuscripts have mag. 3. No authority for mag. 3-4. All manuscripts have mag. 3. No authority for mag. 3-4. B. M. 7475, mag. 2; all others mag. 3. No authority for mag. 3-4. B. M. 7475, mag. 3. } All manuscripts have mag. 3. No authority for mag. 3-4. B. M. 7475, mag. 3.
46 55	DRACO. 3. 23 β 12. 63 ϵ	All manuscripts have mag. 3. No authority for mag. 3-4. All manuscripts have mag. 4. No authority for mag. 4-5. Sûfi has 4-3. Schjellerup gives Ptolemy's mag. 4-3, but the authority is unknown.
57 58 59 62 63 65 69 70	14. 61 σ 15. 52 ν 16. 60 τ 19. 43 φ 20. 27 f 22. 18 g 26. 13 θ 27. 12 ι	All manuscripts have mag. 5. Peters assigns mag. 5-6 to these stars the authority for which is unknown. Sûfi gives mag. 5-4. All manuscripts have mag. 4. No authority for mag. 4-5. B. M. 7475, mag. 4. B. M. 7475, mag. 6. B. M. 7475, mag. 3-2. Most manuscripts have mag. 3. Magnitude 3-2 adopted by Peters is found in Laur. 45 (Gerard of Cremona). Magnitude in B. M. 7475 is doubtful; the scribe gives both 3-2 and 4-3.
71 72 73 74	28. 10 i 29. 11 α 30. 5 κ 31. 1 λ	Bod. 369 and B. M. Reg. 16, mag. 3. All manuscripts give mag. 3. B. M. 7475 gives both 3 and 4. Peters adopted mag. 3-4, which is found only in Sûfi.
80 87	CEPHEUS. 6. 2 θ Inf. 2. 27 δ	B. M. 7475, mag. 3. All manuscripts have mag. 4. No authority for mag. 4-5.
95 108	BOOTES. 8. 51 μ 21. 4 τ	B. M. 7475, mag. 6. B. M. 7475, mag. 3.
111 112 114	COR. BOR. 1. 5 α 2. 3 β 4. 9 π	B. M. 7475 has the singular error of mag. 5-4 for both stars. Ven. 313 and Laur. 48, mag. 5.

PTOLEMY'S CATALOGUE OF STARS.
Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
HERCULES.		
122	4. 7 κ	B. M. 7475, mag. 3; all other manuscripts mag. 4. No authority for mag. 4-5.
123	5. 65 δ	Vat. 1594, Vat. 1038, Venice manuscripts 313, 312, and 310, and Laur. 48, mag. 4.
124	6. 76 λ	B. M. 7475, mag. 3.
128	10. 92 ξ	All the Greek manuscripts have mag. 4, and the Arabs, B. M. 7475, Bod. 369, B. M. Reg. 16, Laur. 45, and Sloane 2795, mag. 4-3, which is adopted.
129	11. 40 ζ	All the Greeks and Baily, mag. 4; B. M. 7475, mag. 4-3; Bod. 369 and B. M. Reg. 16, mag. 3.
130	12. 58 ϵ	All Greek manuscripts and Baily have mag. 5; B. M. 7475, Bod. 369, B. M. Reg. 16, and Laur. 45, mag. 4-3, which is better than mag. 5-6 adopted by Peters, for which no authority is known.
132	14. 61 c	All the Greeks and Baily have mag. 3; the Arabs mag. 5, which is adopted by Peters and accords with the star.
133	15. 67 π	All the Greeks have mag. 4, which is adopted; the Arabs mag. 3.
140	22. 82 y	Bod. 369 and B. M. Reg. 16, mag. 6-5.
142	24. 35 σ	Bod. 369 and B. M. Reg. 16, mag. 4-3.
144	26. 11 φ	} B. M. 7475 makes these stars mag. 6.
145	27. 6 ν	
146	28. 1 χ	
147	29. $\begin{matrix} \nu^1 \\ \nu^2 \end{matrix}$	
LYRA.		
150	2. $\begin{matrix} 4 \epsilon^1 \\ 5 \epsilon^2 \end{matrix}$	} B. M. 7475, mag. 4.
151	3. $\begin{matrix} 6 \zeta^1 \\ 7 \zeta^2 \end{matrix}$	
154	6. 21 θ	
156	8. 9 ν^2	The Greeks and B. M. 7475 have mag. 4; Bod. 369 and B. M. Reg. 16 are the only authorities for mag. 4-5, adopted.
157	9. 14 γ	B. M. 7475, mag. 4; Ven. 302, mag. 4-3.
158	10. 15 λ	B. M. 7475, mag. 3-4. B. M. 7475, mag. 4.
CYGNUS.		
167	9. 1 κ	B. M. 7475, mag. 4.
169	11. 54 λ	Ven. 313 and B. M. 7475, mag. 4.
170	12. 64 ζ	B. M. 7475, mag. 2.
172	14. 62 ξ	B. M. 7475, Laur. 45 (Gerard of Cremona), mag. 4.
173	15. $\begin{matrix} 30 \sigma^1 \\ 31 \sigma^1 \end{matrix}$	} All manuscripts give mag. 4 to these stars. Peters assigns 4-5 to both, but the authority is not known.
174	16. 32 σ^2	
CASSIOPEIA.		
180	3. 24 η	B. M. 7475, mag. 3.
185	8. 33 θ	B. M. 7475, mag. 3.
188	11. 15 κ	B. M. 7475, mag. 4.
PERSEUS.		
192	2. 15 η	Vat. 1038, mag. 3.
196	6. 18 $H \iota$	All manuscripts give mag. 4. Authority unknown for Peters' mag. 4-3.
198	8. 35 σ	B. M. 7475, mag. 3.

Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
PERSEUS—CONT.		
211	21. 58 ϵ	Bod. 369 and B. M. Reg. 16, mag. 5-4.
215	25. 38 \circ	B. M. 7475, mag. 3-2; Bod. 369, mag. 3.
216	26. 44 ζ	Bod. 369, mag. 3-4.
AURIGA.		
228	9. 8 ζ	B. M. 7475, Laur. 45, Sloane 2795, mag. 4-5.
230	11. 23 γ	B. M. 7475, Bod. 369, B. M. Reg. 16, mag. 3-4. This is the same star as No. 400, but the magnitudes given in most cases to the latter do not accord with No. 230.
OPHIUCHUS.		
234	1. 55 α	B. M. 7475 and all manuscripts of Gerard of Cremona, mag. 3.
238	5. 27 κ	B. M. 7475, mag. 3.
239	6. 10 λ	
240	7. 1 δ	All Greek manuscripts give mag. 3; B. M. 7475, mag. 4; Bod. 369 and B. M. Reg. 16, mag. 4-5. No authority is found for 3-4 assigned by Peters.
241	8. 2 ϵ	B. M. 7475, mag. 4-5.
243	10. 64 ν	B. M. 7475, mag. 4.
244	11. 69 τ	B. M. 7475, mag. 4-3.
245	12. 35 η	B. M. 7475, mag. 4.
246	13. 40 ξ	All authorities, Greek and Arabic, agree mag. 4-3. Peters gives 4-5, which is the same as Süfi. In his rough draft of catalogue Peters gives 4-3.
247	14. 36 Λ	Ven. 313, Vat. 1038, mag. 4-3; Bod. 369, mag. 4-5; Laur. 6, mag. 4.
248	15. 42 θ	All authorities give mag. 4-3. No authority known for Peters' mag. 4-5; in rough draft, 4-3.
249	16. 44 b	B. M. 7475, mag. 4-3.
SERPENS.		
271	9. 24 α	B. M. 7475, mag. 4.
273	11. 32 μ	B. M. 7475, mag. 3.
274	12. 3 ν Oph.	B. M. 7475, mag. 4.
AQUILA.		
288	3. 53 α	Ven. 313, Laur. 48, Vat. 1594, Vat. 1038, Ven. 310, mag. 2.
289	4. 59 ξ	Vat. 1038, mag. 3; Bod. 369 and B. M. Reg. 16, mag. 3-2.
290	5. 50 γ	Vat. 1038, mag. 3-4.
DELPHINUS.		
302	2. 5 ι	All Greek authorities, mag. 4-5; the Arabs have mag. 4; Peters gives mag. 4-3.
303	3. 7 κ	Vat. 1038, mag. 4-5.
304	4. 6 β	Vat. 1038, mag. 3.
308	8. 3 η	Bod. 369, mag. 6-7.
PEGASUS.		
319	5. 62 τ	B. M. 7475, mag. 3.
325	11. 42 ζ	Bod. 369, mag. 4.
326	12. 46 ξ	Vat. 1038, mag. 3.
331	17. 8 ϵ	B. M. 7475, mag. 4-3.
334	20. 10 κ	Vat. 1594, Vat. 1038, Ven. 313, Ven. 310, Laur. 48, mag. 4; B. M. 7475, mag. 3-2.

Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
ANDROMEDA.		
345	11. 38 η	Bod. 369, B. M. Reg. 16, Laur. 45, Sloane 2795, mag. 3.
347	13. 37 μ	B. M. 7475, Laur. 45, Sloane 2795, mag. 3.
350	16. φ Pers.	B. M. 7475, Bod. 369, B. M. Reg. 16, Laur. 45, mag. 4-3.
351	17. ν Pers.	Bod. 369, B. M. Reg. 16, mag. 4.
352	18. 50 ν	The Greek manuscripts and B. M. 7475 have mag. 4; Bod. 369 and B. M. Reg. 16, mag. 4-3. No authority found for Peters' mag. 4-5.
TRIANGULUM.		
358	1. 2 α	B. M. 7475, mag. 4.
ARIES.		
375	Inf. 1. 13 α	B. M. 7475, mag. 3.
TAURUS.		
390	11. 54 γ	B. M. 7475, Laur. 45, mag. 4-5.
394	15. 74 ϵ	Bod. 369, Laur. 48, mag. 3.
396	17. 104 m	Vat. 1594, Vat. 1038, Ven. 313, Ven. 310, Laur. 48, Laur. 54, mag. 4.
400	21. 112 β	Laur. 48, Ven. 310, mag. 3-2; Bod. 369, mag. 5; Laur. 45, mag. 4.
404	25. 50 ω^2	Bod. 369, mag. 5.
GEMINI.		
440	17. 24 γ	Baily gives mag. 3.
441	18. 31 ξ	B. M. 7475, Laur. 45, mag. 4-3.
443	Inf. 2. κ Aur.	B. M. 7475, mag. 4.
444	Inf. 3. 36 d	All Greek and Arabic manuscripts have mag. 5. No authority for Peters' mag. 5-6.
445	Inf. 4. 85	B. M. 7475, mag. 3-4.
CANCER.		
453	5. 47 δ	B. M. 7475, mag. 4.
LEO.		
465	4. 17 ϵ	B. M. 7475, Laur. 45, mag. 3.
466	5. 36 ζ	Bod. 369, mag. 2.
480	19. 60 b	All Arabs and Laur. 6 have mag. 5; the Greeks and Baily, mag. 6.
481	20. 68 δ	B. M. 7475, Laur. 45, mag. 2.
483	22. 70 θ	Bod. 369, mag. 5.
487	26. 91 ν	B. M. 7475, mag. 4.
494	See Notes on the Catalogue of Stars.
VIRGO.		
506	10. 43 δ	B. M. 7475, mag. 4.
509	13. 47 ϵ	Paris 2389, Vat. 1594, Vat. 1038, Ven. 313, Ven. 312, Ven. 302 Ven. 310, and Laur. 48, have mag. 5-4, and this has been adopted by Peters. B. M. 7475 and Sloane 2795, and Laur. 45, mag. 3; Bod. 369 and B. M. Reg. 16, mag. 3-2. Süfi describes the star as of mag. 3 and adds: "Ptolémée la dit des moindres; that means mag. 3-4. Manitius has adopted mag. 3-2, which is more correct for ϵ Virginis than 5-4. Baily gives mag. 5.
510	14. 67 α	All authorities give mag. 1. Peters has adopted mag. 1-2 as given by Süfi.
511	15. 79 ζ	Bod. 369, mag. 3-2.
514	18. 82 m	Laur. 6, mag. 4-3.
520	24. 105 φ	All authorities have mag. 4. Peters adopts mag. 4-5 as in Süfi.

Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
VIRGO—continued.		
522	26. 107 μ	All the Greeks have mag. 3; B. M. 7475, Bod. 369, Laur. 45 have mag. 4; Manitius gives mag. 4; Peters adopts mag. 3-4; Sûfi has mag. 4-3.
526	Inf. 4. 53	B. M. 7475, mag. 5.
SCORPIUS.		
558	13. $\left\{ \begin{array}{l} \mu^1 \\ \mu^2 \end{array} \right.$	B. M. 7475, Laur. 45, Sloane 2795, mag. 4.
568	Inf. 2. 45 d Oph.	All authorities agree. Manitius has mag. 5.
SAGITTARIUS.		
574	5. $\left\{ \begin{array}{l} \mu^1 \\ \mu^2 \end{array} \right.$	B. M. 7475, mag. 3.
576	7. 27 φ	Paris 2389, Vat. 1594, Ven. 313, and Manitius have mag. 4; B. M. 7475, mag. 3-2; Bod. 369, and B. M. Reg. 16, mag. 4-3. No authority for Peters' mag. 4-5.
588	19. $\left\{ \begin{array}{l} 51 h^1 \\ 52 h^2 \end{array} \right.$	B. M. 7475, Laur. 45, and Sloane 2795, mag. 5.
594	25. XVIII. 17	Laur. 48, mag. 3-4.
CAPRICORNUS.		
602	2. 8 ν	Bod. 369, mag. 5.
607	7. 11 ρ	Bod. 369, mag. 5.
608	8. 7 σ	Ven. 312, mag. 6.
611	11. 16 ψ	B. M. 7475, Laur. 45, Sloane 2795, mag. 6.
AQUARIUS.		
632	4. 22 β	Gerard of Cremona in three manuscripts, mag. 2.
673	Inf. 3. 7 Ceti.	Ven. 313, Ven. 312, Vat. 1038, Laur. 45, Sloane 2795, mag. 4.
PISCES.		
674	1. 4 β	Vat. 1594, Vat. 1038, Ven. 313, Ven. 312, Ven. 310, Laur. 48, and Manitius, mag. 4. The mag. 4-3, adopted, is from Paris 2389, Ven. 302, Laur. 6, Laur. 45, B. M. 7475, and Bod. 369.
700	27. 67 k	Ven. 313, mag. 5.
CETUS.		
726	15. 45 θ	Laur. 45, Sloane 2795, mag. 2.
727	16. 31 η	Baily gives mag. 5 for which Grynæus is the only authority.
730	19. 17 φ^1	Vat. 1594, Vat. 1038, Ven. 313, Ven. 310, Laur. 48, Laur. 45, and B. M. 7475, mag. 5.
731	20. O. 161	All manuscripts have mag. 5, except Paris 2389, Ven. 302, Ven. 312, B. M. Reg. 16, and Bod. 369, mag. 5-4.
732	21. 8 ι	Vat. 1594, Vat. 1038, Ven. 313, Ven. 310, Laur. 48, Laur. 45, have mag. 3; Laur. 6, 5-6; B. M. 7475, mag. 4.
733	22. 16 β	All Greek and Latin authorities have mag. 3. Bod. 369 and B. M. Reg. 16 have mag. 3-4; and B. M. 7475, mag. 3-2, which is the only authority found for the magnitude adopted by Peters.

Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
ORION.		
736	3. 24 γ	All authorities have mag. 2, except the Arabs, Laur. 45, B. M. 7475, Bod. 369, and B. M. Reg. 16, which have 2-1.
742	9. 72 f^2	B. M. 7475, mag. 4.
763	30. $\left\{ \begin{array}{l} 42 c \\ 45 c \end{array} \right.$	Ven. 312 and Laur. 45, mag. 4-5.
764	31. $\left\{ \begin{array}{l} 41 \theta^1 \\ 43 \theta^2 \end{array} \right.$	Paris 2389, Ven. 313, Ven. 312, Vat. 1038, Laur. 6, and B. M. 7475, mag. 3.
765	32. 44 ι	Paris 2389, Ven. 312, Vat. 1038, and Laur. 6, mag. 3-4.
769	36. 20 τ	B. M. 7475, mag. 4.
771	38. 53 κ	B. M. 7475 and Laur. 45, mag. 3.
ERIDANUS.		
772	1. 69 λ	Vat. 1038, Laur. 45, and B. M. 7475, mag. 4.
790	19. 2 τ^2	B. M. 7475, mag. 3.
In B. M. 7475 the magnitudes of these stars are omitted, but in the place of each magnitude is written the Arabic letter Kaf. This might be taken for the initial of the word Kabir, which is the Arabic for $\mu\epsilon\iota\zeta\omega\nu$, but in this manuscript the $\mu\epsilon\iota\zeta\omega\nu$ and $\epsilon\lambda\acute{\alpha}\sigma\sigma\omega\nu$ are invariably represented by the letters Mim and Lam. See description of B. M. 7475.		
802	31. III 202	See Notes to the Catalogue of Stars.
803	32. III 189	
804	33. III 149	
805	34. θ	
CANIS MAJOR.		
818	1. 9 a	Bod. 3374, mag. 4.
819	2. 14 θ	B. M. 7475, mag. 3.
824	7. 8 ν^2	The Greeks and Baily give mag. 6; all the Arabs and Manitius have mag. 5.
828	11. 5 ξ^2	B. M. 7475, mag. 4.
830	13. 16 o^1	Vat. 1038, mag. 5-6.
831	14. 25 δ	Vat. 1038 and B. M. 7475, mag. 3.
835	18. 31 η	B. M. 7475, mag. 4-5.
ARGO NAVIS.		
854	6. VII 175	B. M. 7475 and manuscripts of Gerard of Cremona, mag. 4.
860	12. π Pup.	Laur. 48, mag. 4.
861	13. f Pup.	
863	15. c Pup.	All authorities, mag. 4. No authority found for Peters' mag. 4-5.
874	26. b Vel.	Vat. 1038, Laur. 45, and B. M. 7475, mag. 4.
876	28. a Pyx.	All authorities (except Gerard of Cremona, mag. 4) agree in mag. 3.
885	37. o Pup.	No authority is found for Peters' mag. 3-4.
888	40. κ Vel.	Mag. 2 adopted by Peters and Baily, is confirmed by Paris 2389, Vat. 1594 and Ven. 313, and the printed editions of Grynæus and Trapezuntius; all others, including Manitius, mag. 3.
893	45. τ Pup.	All authorities have mag. 3. No authority found for Peters' mag. 3-4. Aboul Hhassan, who derived his magnitudes from Sûfi, gives 4-3.
HYDRA.		
909	16. 42 μ	B. M. 7475, mag. 2.
912	19. 11 β Crat.	All authorities have mag. 3. Peters' mag. 3-4 is the same as Sûfi. The Arabs have mag. 4; Sûfi also.

Notes on the Star Magnitudes—continued.

Baily's No.	Star.	Notes.
CRATER.		
923	3. 12 δ	All authorities, mag. 4. No authority found for Peters' mag. 4-5. B. M. 7475, mag. 4-3. Vat. 1038, Laur. 48, Laur. 45, and B. M. 7475, mag. 4; Bod. 369 and B. M. Reg. 16, mag. 4-3.
925	5. 14 ε	
926	6. 30 η	
CENTAURUS.		
937	3. 1 i	Paris 2389, Vat. 1594, Ven. 313, Bod. 369, B. M. Reg. 16, and Manilius, mag. 4-3; B. M. 7475, mag. 5-4. No authority found for Peters' mag. 4-5. In rough draft of catalogue he gives mag. 4-3.
945	11. b	Ven. 313, mag. 4-5. Bod. 369, mag. 4-3.
950	16. η	
951	17. κ	Vat. 1594, Vat. 1038, Ven. 310, Laur. 48, mag. 4-3. Vat. 1594, Ven. 302, Vat. 1038, Ven. 310, and Laur. 48, mag. 3.
952	18. ζ	
967	33. δ Crucis	Grynæus transposes the magnitudes of these stars which Baily has erroneously copied.
968	34. α Crucis	
LUPUS.		
985	14. η	Laur. 45 and B. M. 7475, mag. 4-3.
COR. AUST.		
1006	9. ε	B. M. 7475, mag. 4.
PIS. AUST.		
1015	5. 18 ε	Laur. 45 and B. M. 7475, mag. 4. Laur. 45, Bod. 369 and B. M. Reg. 16, mag. 3. Baily gives mag. 3 from Grynæus; no other authority known.
1025	Inf. 3. ε Mic.	
1028	Inf. 6. 24 A Cap.	

TABLE IX.

Collations of Manuscripts.

The Table of Collations gives the variants in longitude and latitude from Baily's Ptolemy in the following 26 manuscripts of the *Almagest*:

Greek: Paris 2389, 2390, 2391, and 2394; Venice 302, 303, 310, 311, 312, and 313; Vatican 1594, 1038, and Reg. 90; Laurentian 1, 47, and 48; Bodleian 3374, and Vienna 14.

Latin: Laurentian 6 and 45; Vienna 24, and British Museum Sloane 2795.

Arabic: British Museum 7475 and Reg. 16; Bodleian 369, and Laurentian 156.

For the purpose of comparison, readings agreeing with Baily are given in doubtful cases, and in those which instance peculiar mistakes of the copyist.

All Baily's readings which differ from the Catalogue are given in the Notes to the Catalogue.